

High Spin States in Nuclei: Exotic Quantal Rotation

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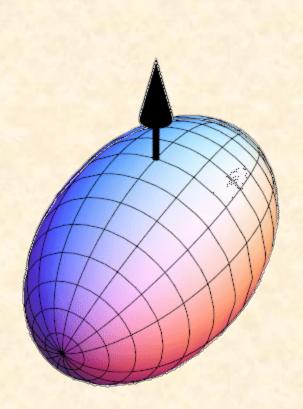
Supported in part by the National Science Foundation

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- "normal" collective rotation principal axis; prolate deformation
- magnetic rotation
 planar tilted axis; weak deformation
- chiral rotation
 aplanar tilted axis; triaxial shape
- tidal waves
 rotating condensate of phonons

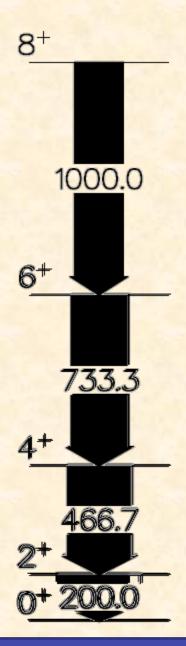




E2 transitions

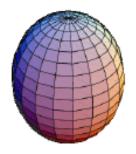
$$E_I = (\hbar^2/2J) I(I+1); E_Y = [](4I-2)$$

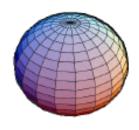
 $\Delta E \gamma = Constant \rightarrow$ "picket fence"

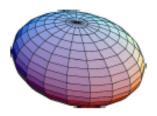




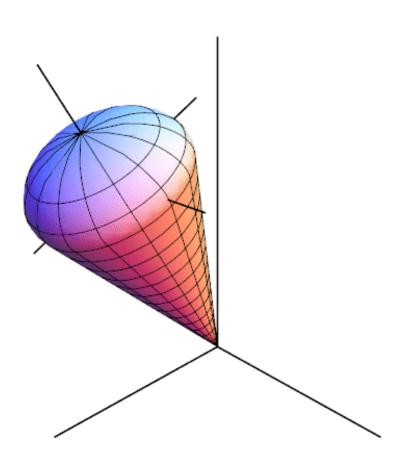
Nuclear shapes: prolate, oblate and triaxial



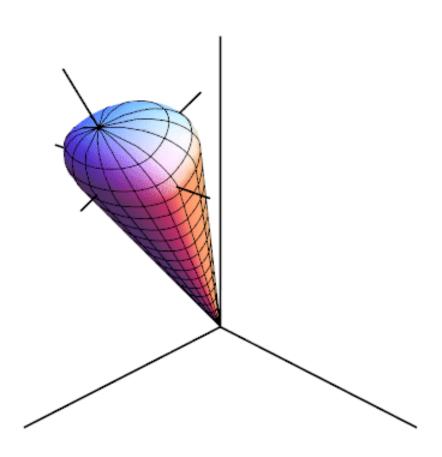




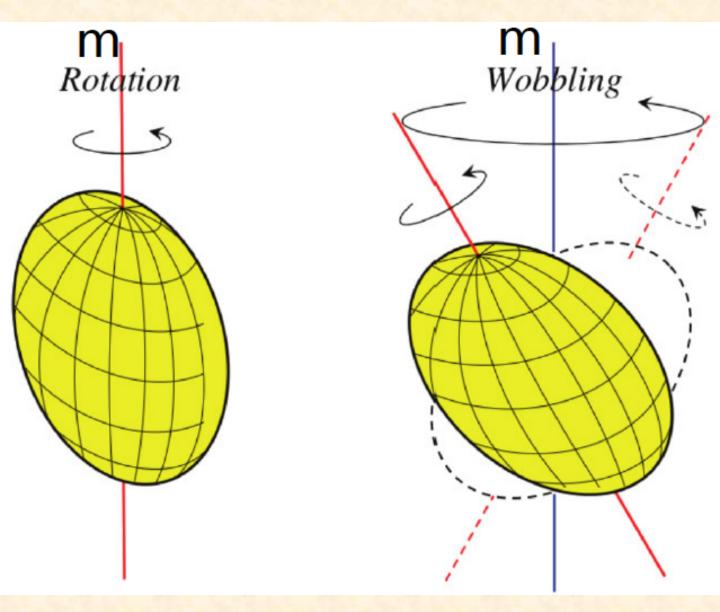




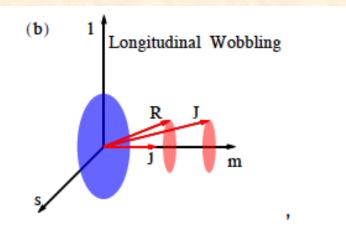


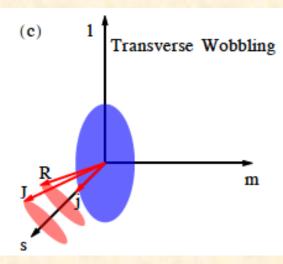




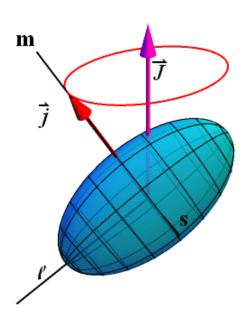




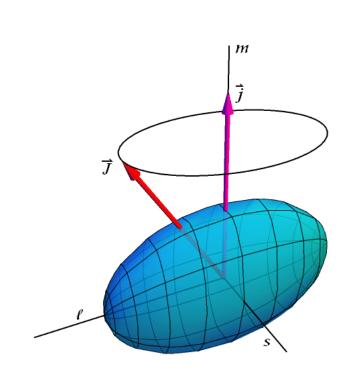




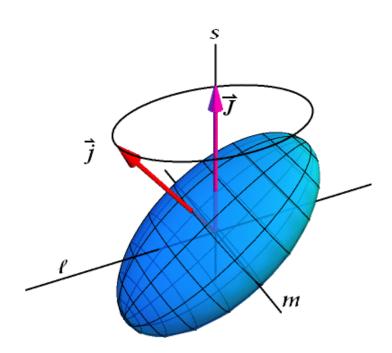




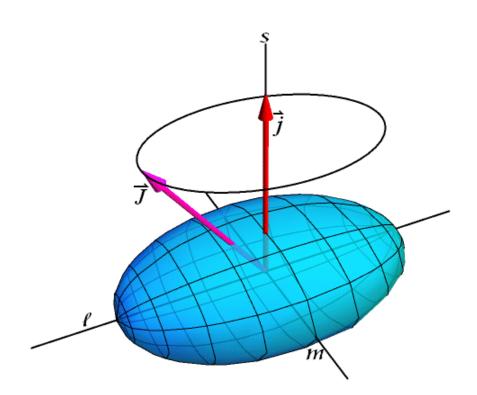












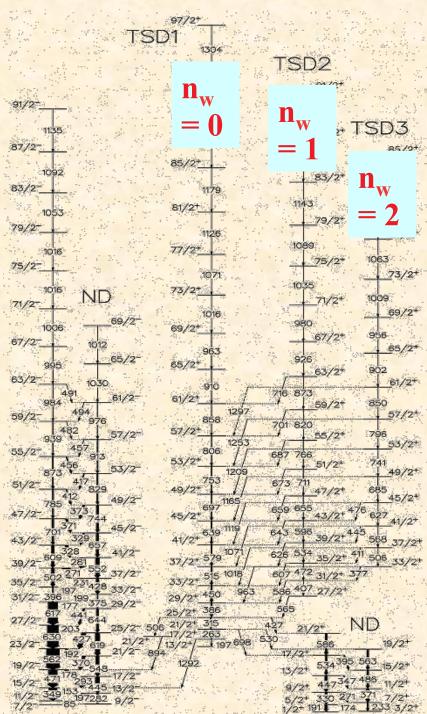


Wobbling bands (TSD) are generally considered as one of the best signatures of nuclear triaxiality. Another is, of course, chirality.

Triaxiality in nuclei had been a longstanding prediction of theory, but had proved very difficult to establish experimentally.

The best example of wobbling has been seen in the Lu nuclei.





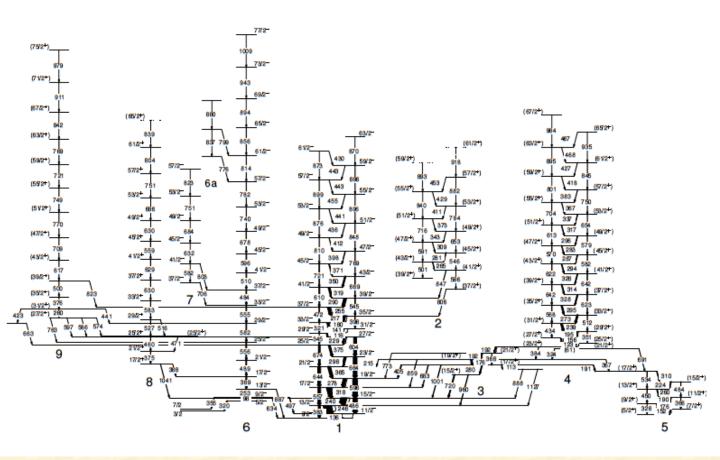


"Wobbler Bands"

- **Rotational bands corresponding to** $n_{\omega} = 0, 1, 2, \dots$
- **Transitions from** $n_{\omega+1} \rightarrow n_{\omega}$ ["one way" and $\Delta n_{\omega} = +1$]
- **❖** Interband transitions are $\Delta J = 1$, *E*2

Observed only in ^{161,163,165,167}Lu and ¹⁶⁷Ta





Nine decay sequences are now known in 169 Re including the $i_{13/2}$ band that is the basis for wobbling structures in nearby nuclei.

Wobbling could not be identified in this nucleus as the $i_{13/2}$ band was relatively weakly populated due to its location high in energy relative to the yrast structures.



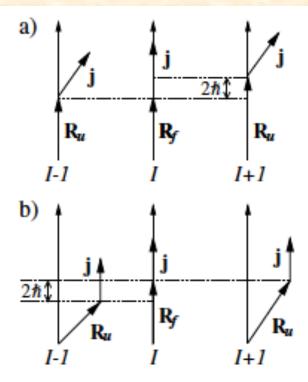


FIG. 4. Schematic coupling scheme of the particle and core angular momenta in the favored (I) and unfavored $(I \pm 1)$ states for (a) the cranking regime and (b) the wobbling mode $(n_w = 1)$. The total angular momentum is I = R + j, where the angular momentum of the collective rotation of the core is expressed by R. The vertical axis shown (x axis) is the axis of the largest moment of inertia of the core, about which collective rotation is energetically cheapest. For $n_w > 1$ the angle between the x axis and R gets larger.



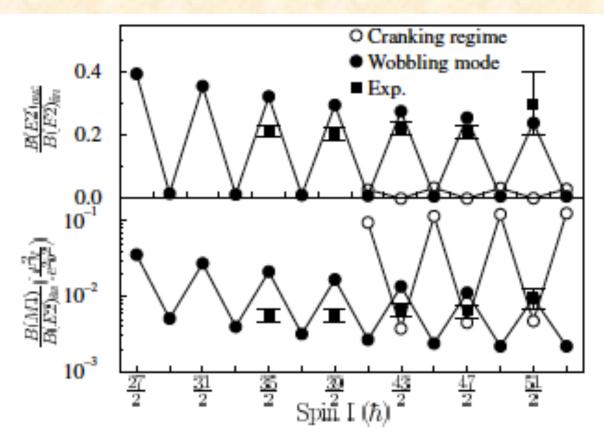
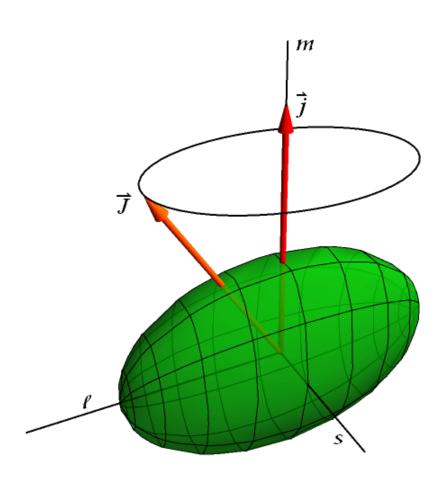


FIG. 5. Experimental and calculated electromagnetic properties of the connecting transitions.







Wobbling frequency, defined by:

$$E_{\text{wobb}} = E (I, n_{\omega} = 1) - [E (I+1, n_{\omega} = 0) + E (I-1, n_{\omega} = 0)]/2$$

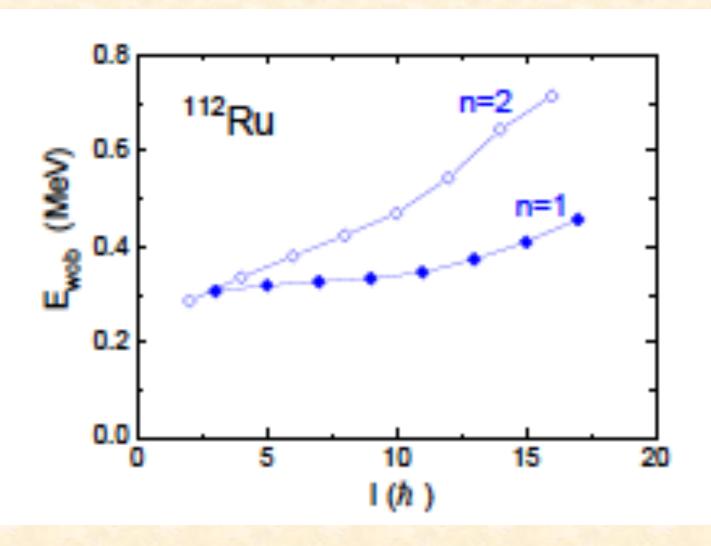
$$\hbar \omega_w = \frac{j}{J_2} \left[\left(1 + \frac{J}{i} \left(\frac{J_3}{J_1} - 1 \right) \right) \left(1 + \frac{J}{i} \left(\frac{J_3}{J_2} - 1 \right) \right) \right]^{1/2}$$

"Longitudinal " wobbler:

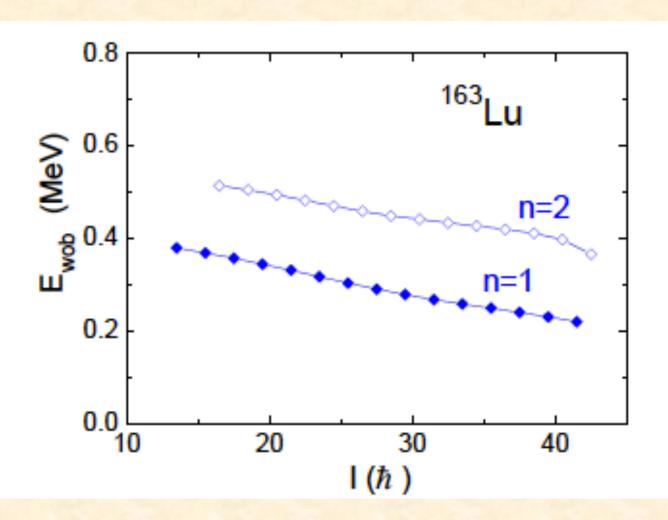
Odd-particle aligned with the axis with maximum moment of inertia (the "medium axis)

$$\mathfrak{I}_3 > \mathfrak{I}_2$$
; $\mathfrak{I}_3 > \mathfrak{I}_1$
 $\rightarrow E_{\omega}$ increases with J



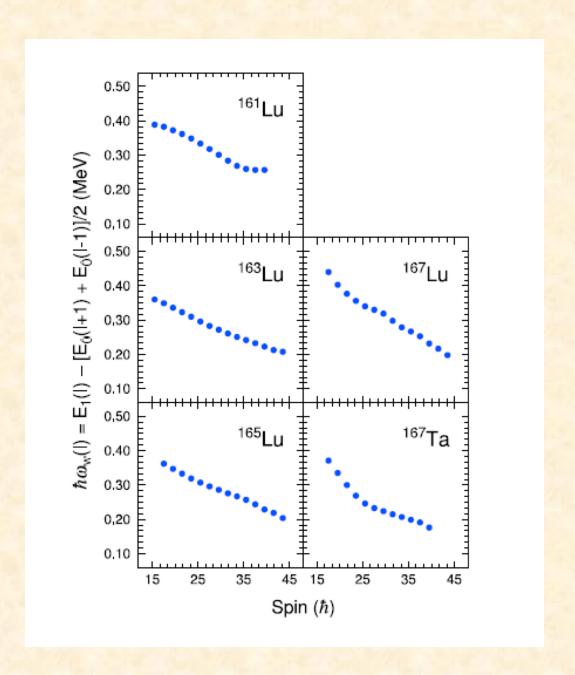






"Standard" wobbler would have increasing E_{wobb}!







Wobbling frequency, defined by:

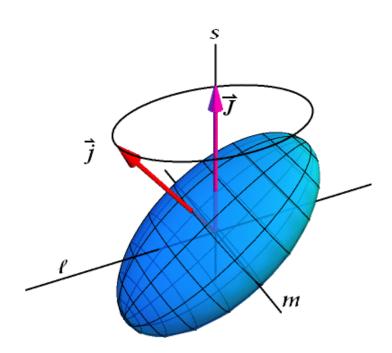
$$E_{\text{wobb}} = E (I, n_{\omega} = 1) - [E (I+1, n_{\omega} = 0) + E (I-1, n_{\omega} = 0)]/2$$

$$\frac{j}{J_3} \left[\left(1 + \frac{J}{i} \left(\frac{J_3}{J_1} - 1 \right) \right) \left(1 + \frac{J}{i} \left(\frac{J_3}{J_2} - 1 \right) \right) \right]^{1/2}$$

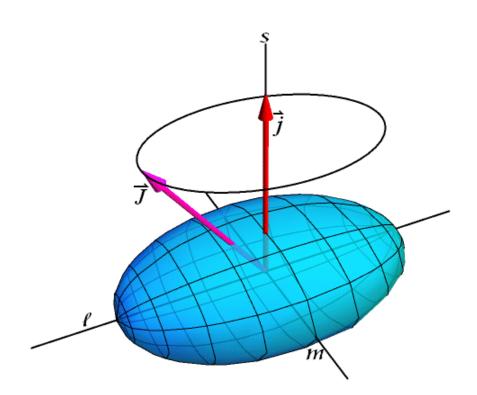
"Transverse" wobbler: Odd-particle aligned with the "small" axis

$$\mathfrak{I}_3 < \mathfrak{I}_2$$
; $\mathfrak{I}_3 > \mathfrak{I}_1$
 $\rightarrow E_{\omega}$ decreases with J
reaching 0 at $J_c = j \mathfrak{I}_2/(\mathfrak{I}_2 - \mathfrak{I}_3)$

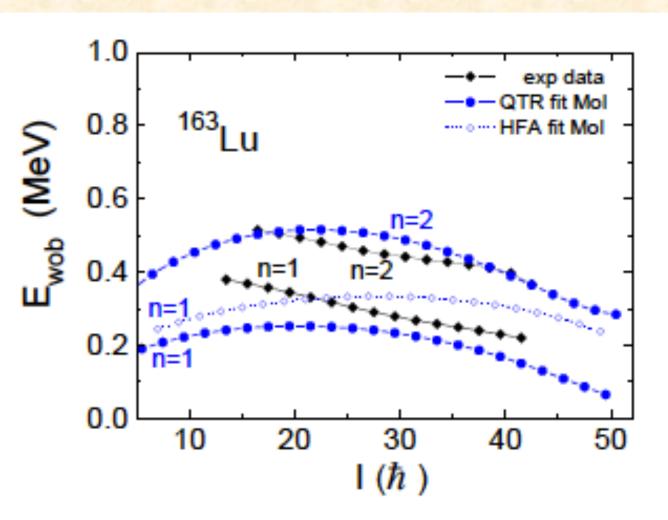




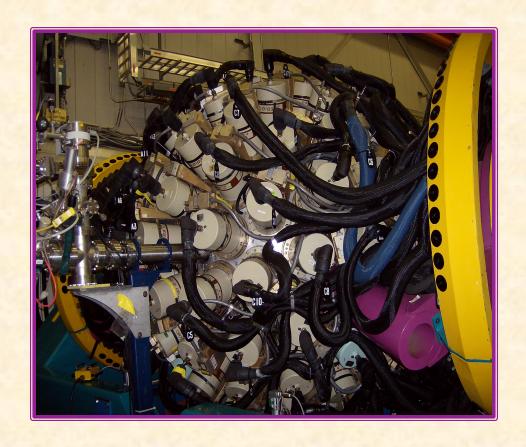






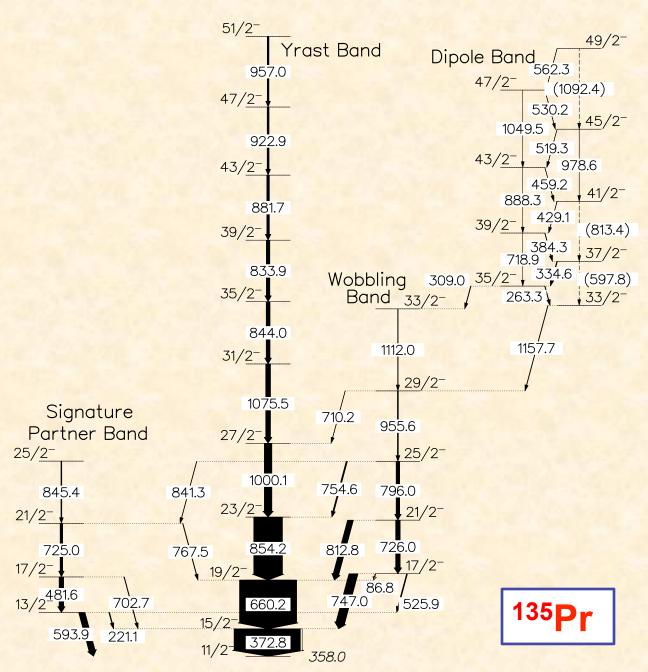






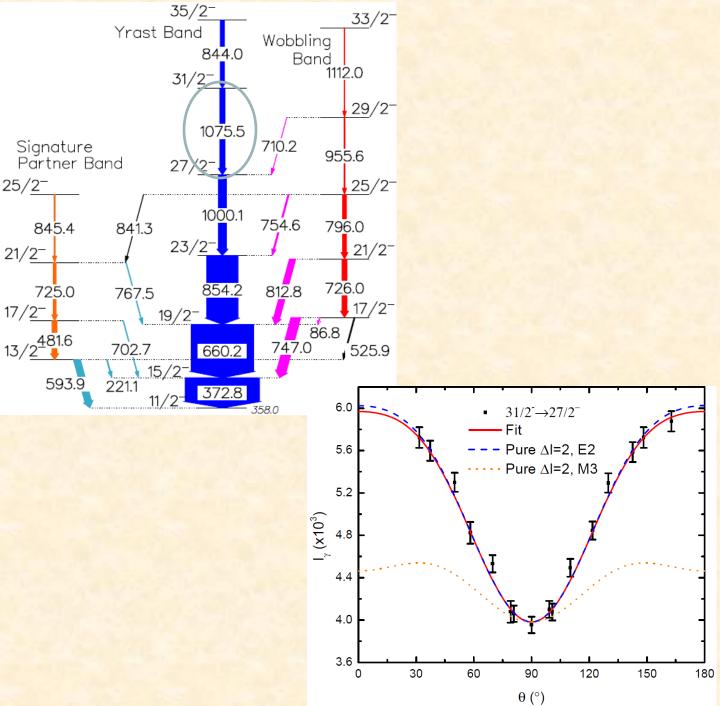
123Sb (16O, 4n)135Pr @ 80 MeV Gammasphere at ATLAS (100 CSGe detectors) γ-γ-γ coincidences angular correlations





J. T. Matta et al., Phys. Rev. Lett. 114, 082501 (2015)







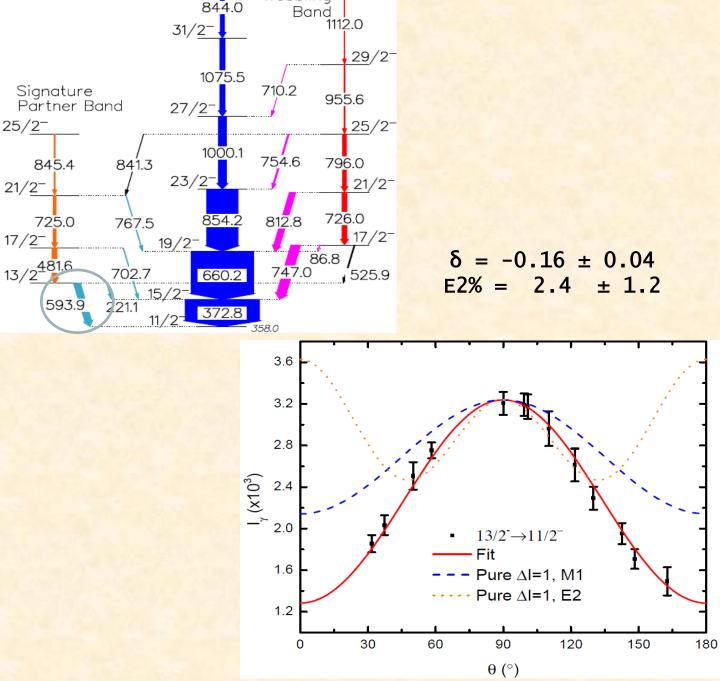
Wobbling

 $33/2^{-}$

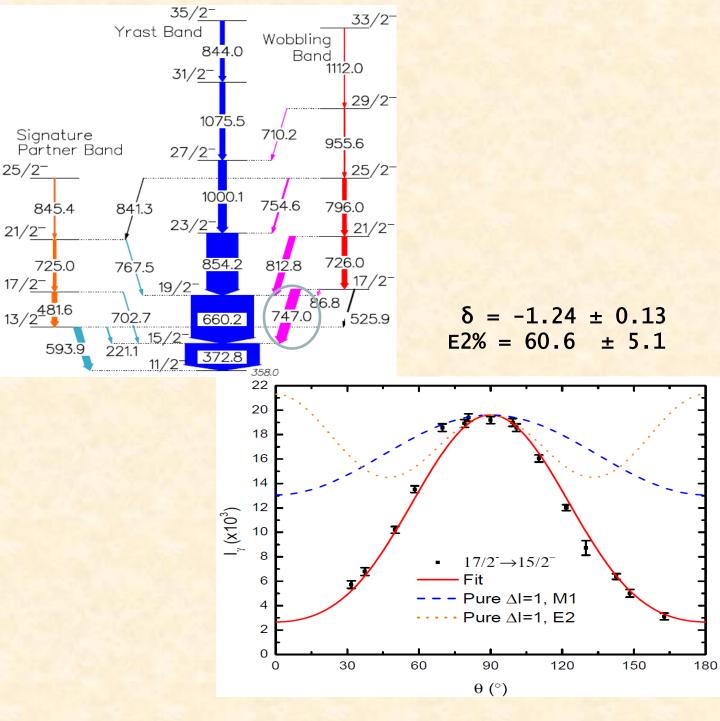
 $35/2^{-}$

844.0

Yrast Band



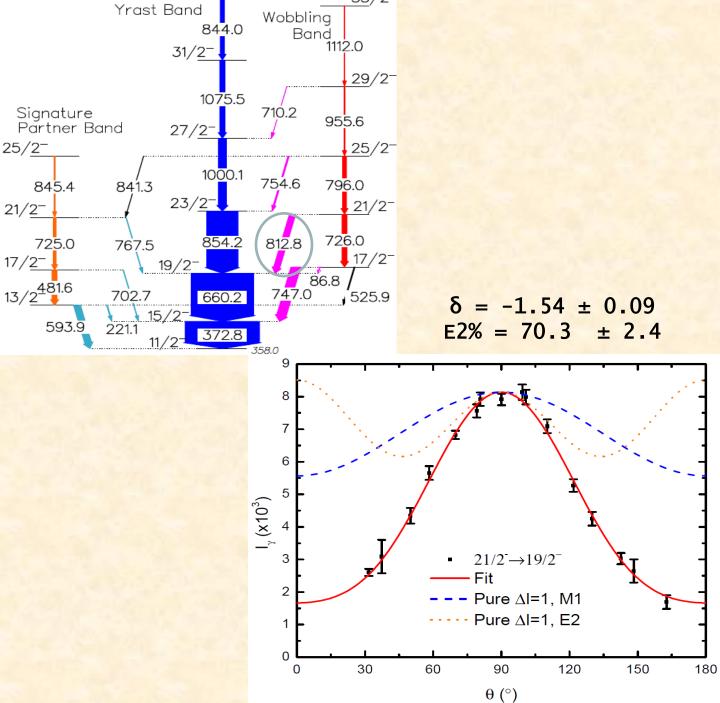






 $33/2^{-}$

 $35/2^{-}$

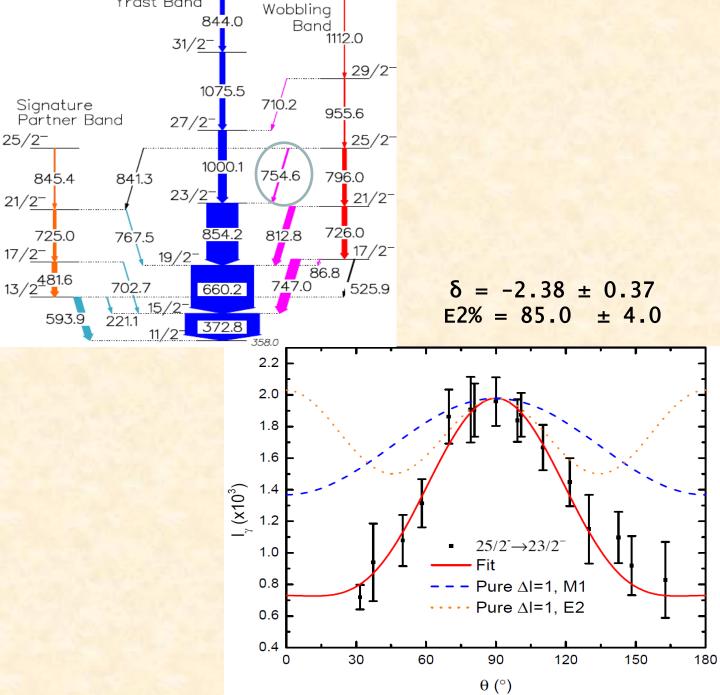




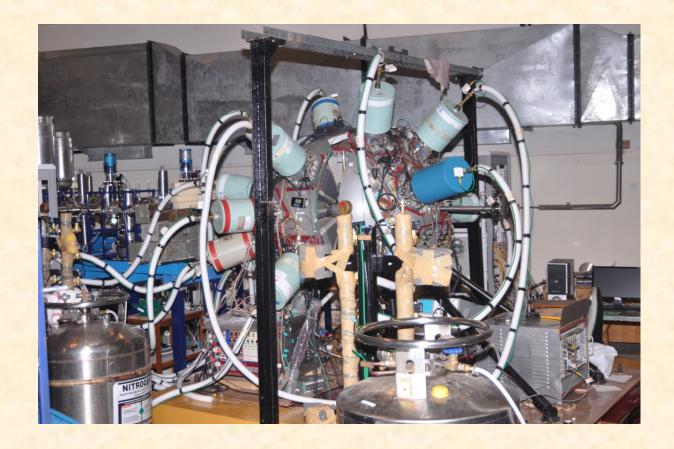
 $33/2^{-}$

 $35/2^{-}$

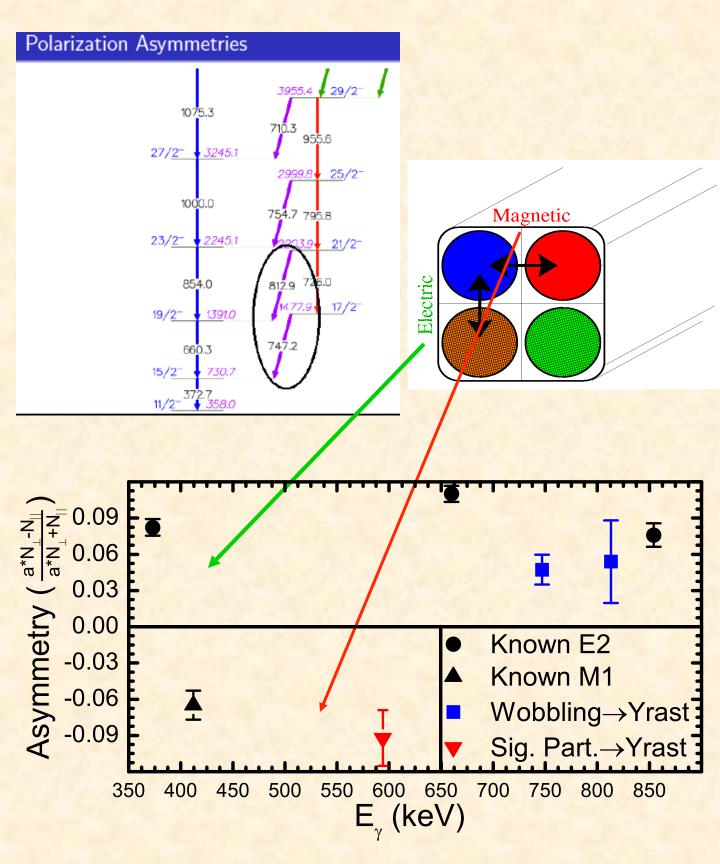
Yrast Band



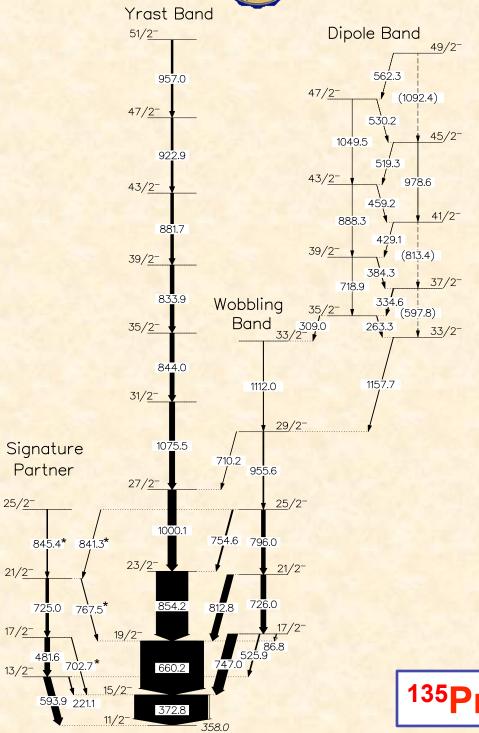




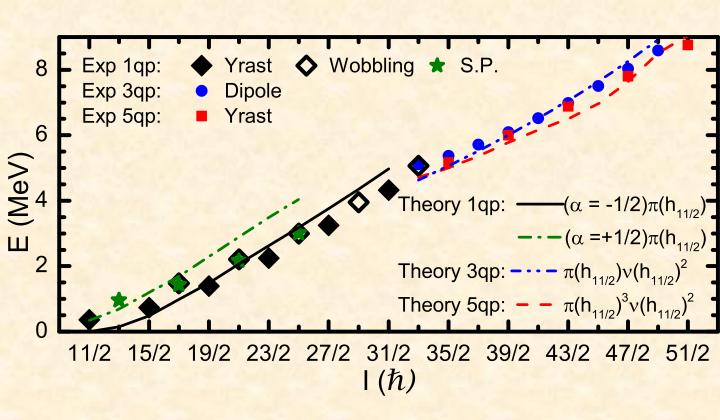
INGA @ TIFR 20 CS "clover" detectors polarization measurements



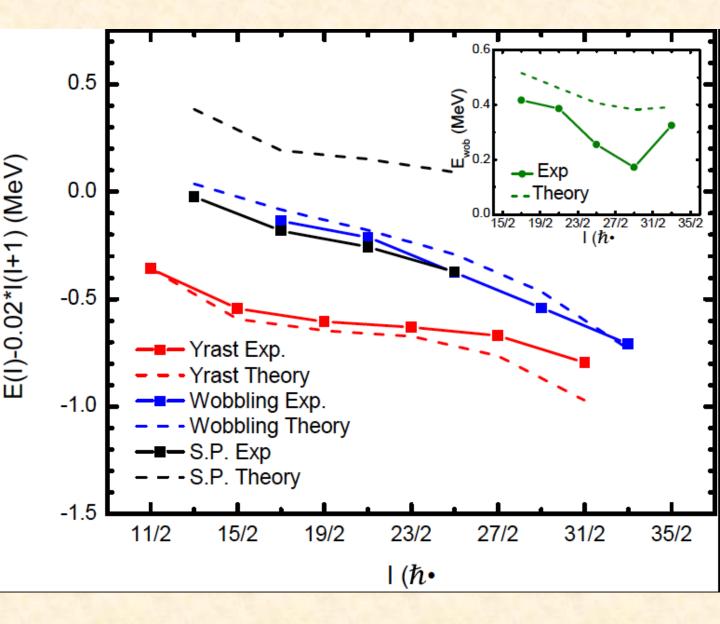














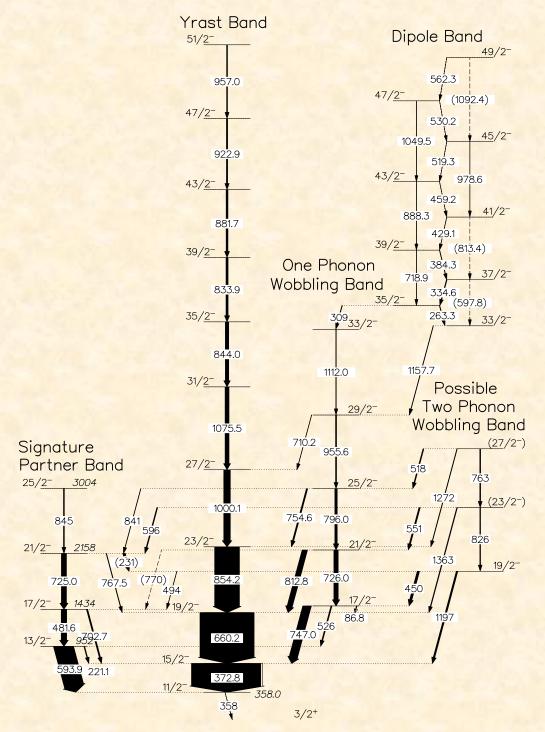
✓ Measurements of level energies, angular distributions, and polarizations of the associated γ rays, have established a "wobbler" sequence in ¹³⁵Pr.

First observation of wobbling in any nuclei away from A~160 region.

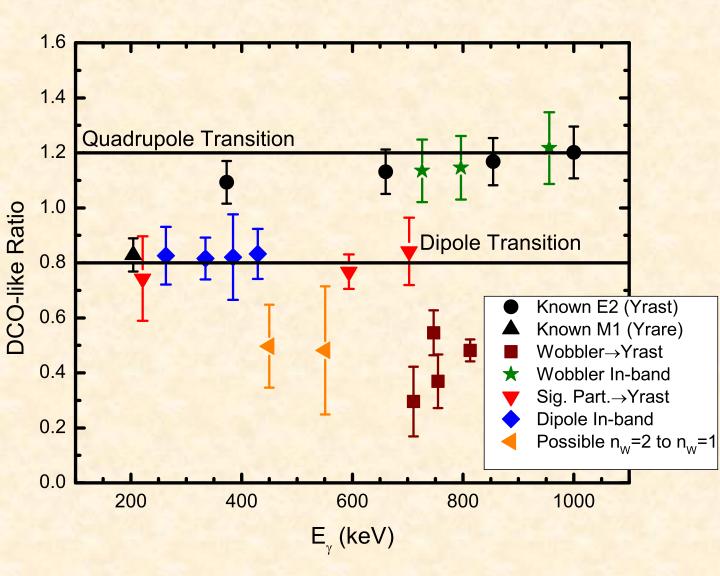
- ✓ Comparison with calculations in QTR model establishes the observed structure as corresponding to a "transverse wobbler"
- ✓ The transmutation of the transverse wobbler into a longitudinal wobbler and then to a magnetic rotation structure is observed in line with theoretical predictions.

Clear indications of gradual change of the rotational axis from "short" into a planar geometry akin to magnetic rotation.

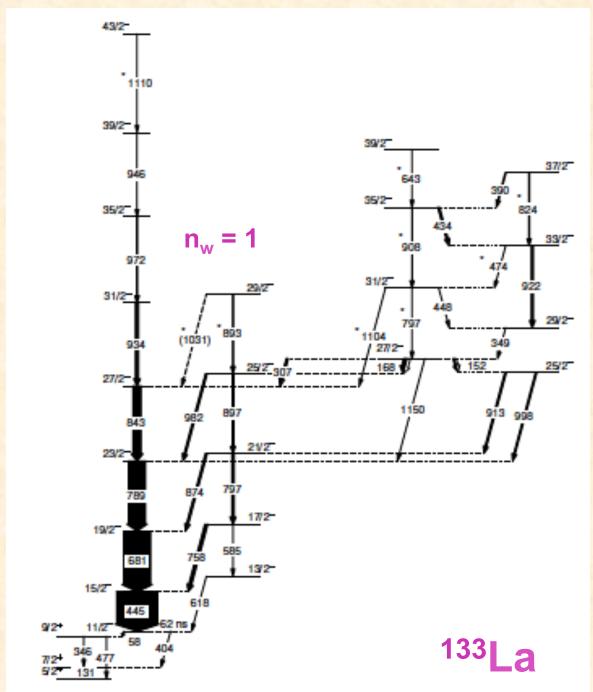




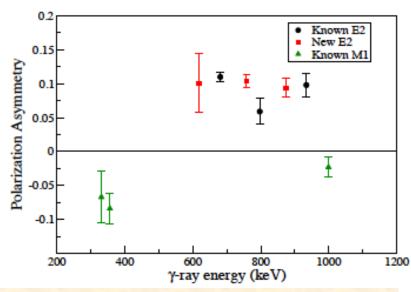




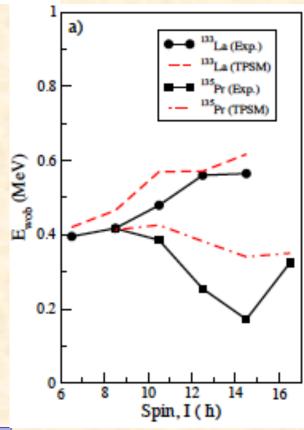








 133 La $n_{\omega} = 1$ band





Octupole Condensation

Strong octupole correlations have been observed in nuclei in the A~230 region.

Such octupole correlations come from long-range interactions between valence nucleons occupying states with $\Delta j = \Delta l = 3$



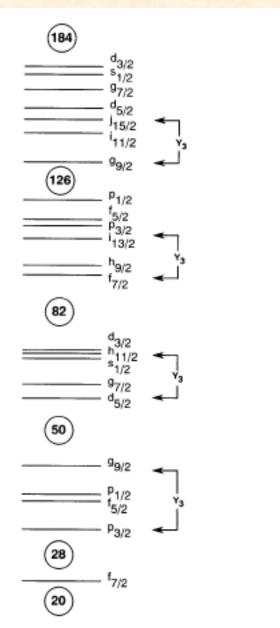


FIG. 4. Nuclear spherical single-particle levels. The most important octupole couplings are indicated.



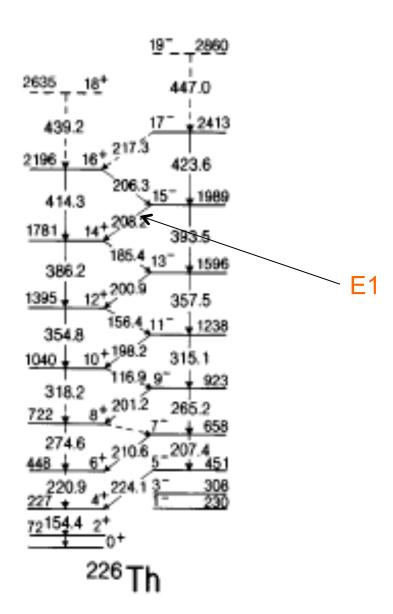


FIG. 17. Level scheme of ²²⁶Th, taken from Schüler et al. (1986; see also Ackermann et al., 1993). The level and transition energies are in keV.



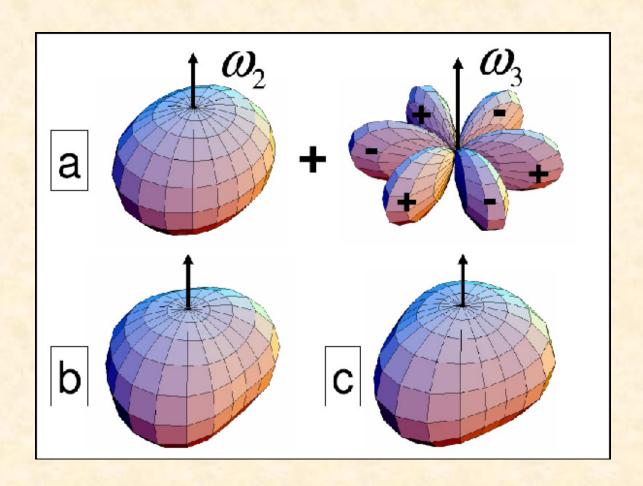
A possible explanation for these correlations has been proposed by Frauendorf in terms of "octupole condensation".

The quadrupole-deformed nucleus is assumed to be a rigid rotor and the octupole vibration is harmonic with a frequency Ω_3 .

No interaction between the octupole phonons and the deformed potential of the nucleus.

The superposition of these two modes resembles an octupole deformed nucleus and the running of the wave gives the appearance of an octupole nucleus rotating or vibrating.

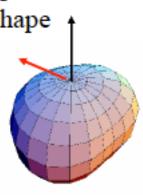




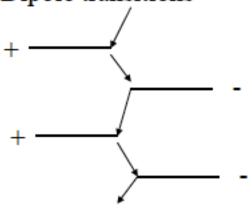
S. Frauendorf, Phys. Rev C **77**, 021304(R) (2008)



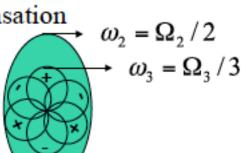
Strong coupling rotating heart shape

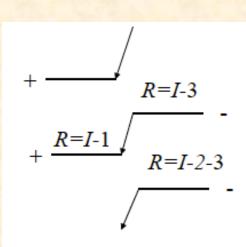


Dipole transitions



Weak coupling phonon condensation







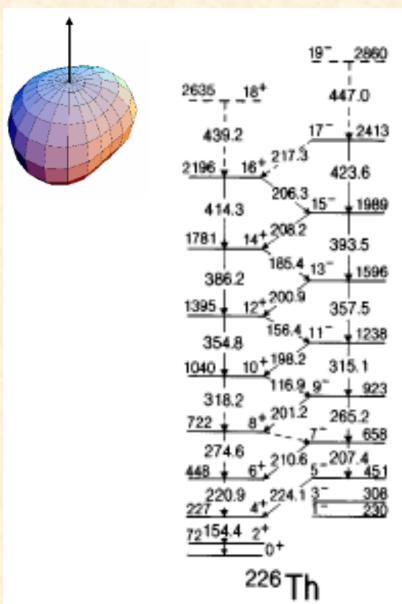
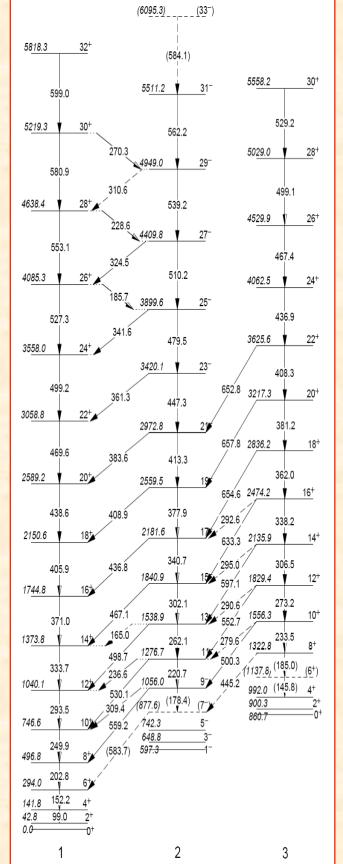
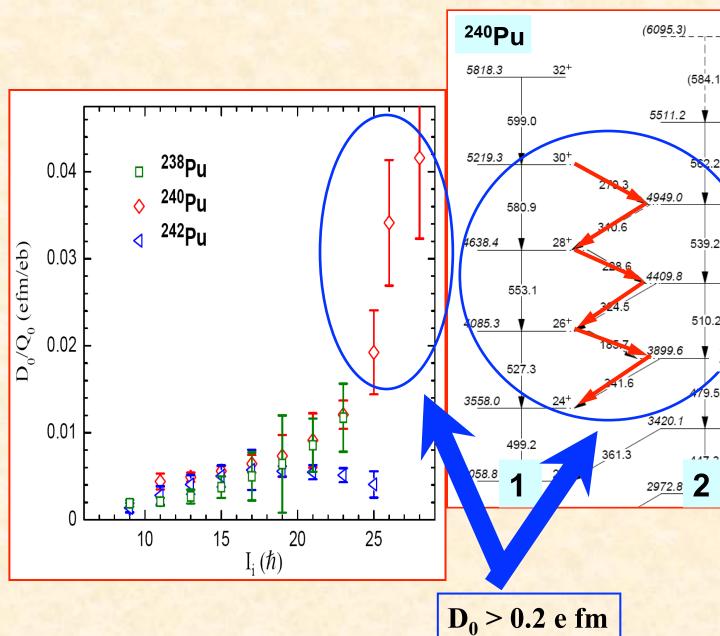


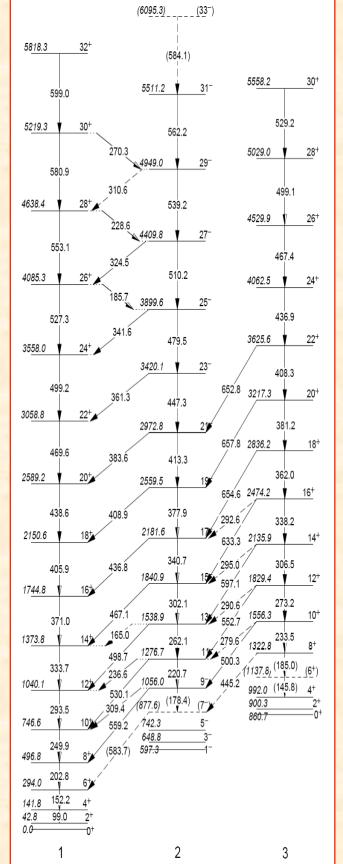
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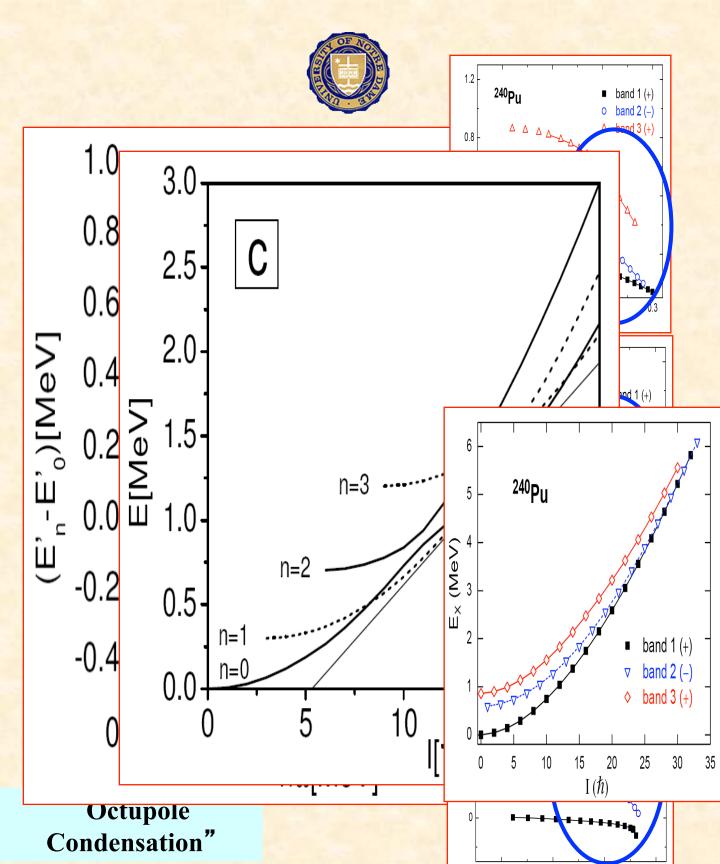
240Pu







240Pu





- ✓ Strong octupole correlations have been observed in the nucleus ²⁴⁰Pu.
- ▼ The data appears to agree well with the predictions of the new "octupole condensation" picture.
- ✓ Another good test case would be ²³⁰Th.

X. Wang et al., Phys. Rev. Lett. 102, 122501 (2009)



Excellent resources (if you are so inclined):

- J. A. Cerny, Editor, Nuclear Reactions and Spectroscopy, Academic Press, 1974.
- H. Morinaga and T. Yamazaki, In-beam Gammaray Spectroscopy, North Holland, 1976.
- H. Ejiri and M.A.J. de Voigt, Gamma-ray and Electron Spectroscopy in Nuclear Physics, Clarendon Press, 1989
- S. Frauendorf, Rev. Mod. Phys. 73, 463 (2001).



ありがとう ध-य वा द Thanks!











The Question Kitten







