

Hypernuclear Structure and Hyperon Star Properties with Relativistic Density Functional Theory

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Hypernuclear systems and neutron stars offer complementary environments for exploring baryon interactions across a wide range of densities. This report investigates Ξ^- hypernuclear structure and the equation of state (EOS) of hyperon-rich matter using relativistic density functional theory. The structure of selected light Ξ^- hypernuclei, such as $^{15}_{\Xi^-}\text{C}$ and $^{13}_{\Xi^-}\text{B}$, are described within the density-dependent relativistic mean-field (DDRMF) framework, where the in-medium behavior of meson-hyperon couplings and the role of the isovector scalar δ meson are analyzed in detail. To extend the constraints toward supranuclear densities, experimental information from hypernuclei is incorporated into a Bayesian inference scheme. A statistically significant linear correlation between scalar and vector ΞN coupling strengths is established and used as a nuclear-physics prior. When combined with multimessenger astrophysical observations, including neutron star masses, radii, and tidal deformabilities, the resulting posterior distribution enables a more constrained and realistic description of hyperonic matter. This approach improves the stiffness of the EOS and supports the existence of $2M_{\odot}$ hyperon stars, offering new insights into the hyperon puzzle from a joint nuclear and astrophysical perspective.

Research field of your presentation

Theoretical Low-energy nuclear physics

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