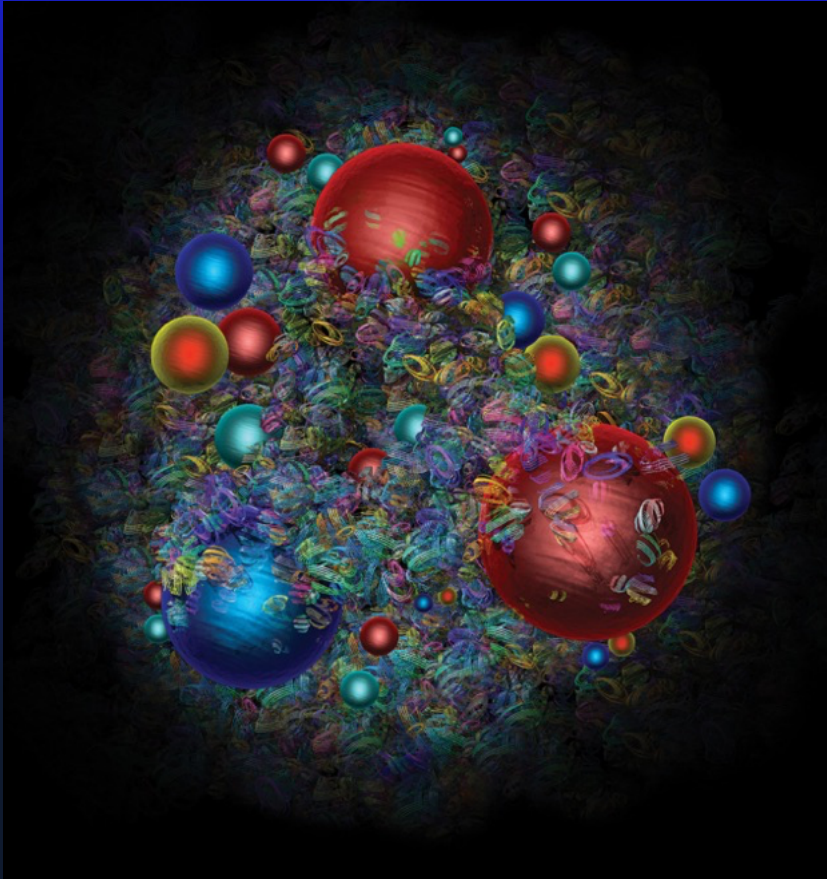


The Electron-Ion Collider: the ultimate electron microscope



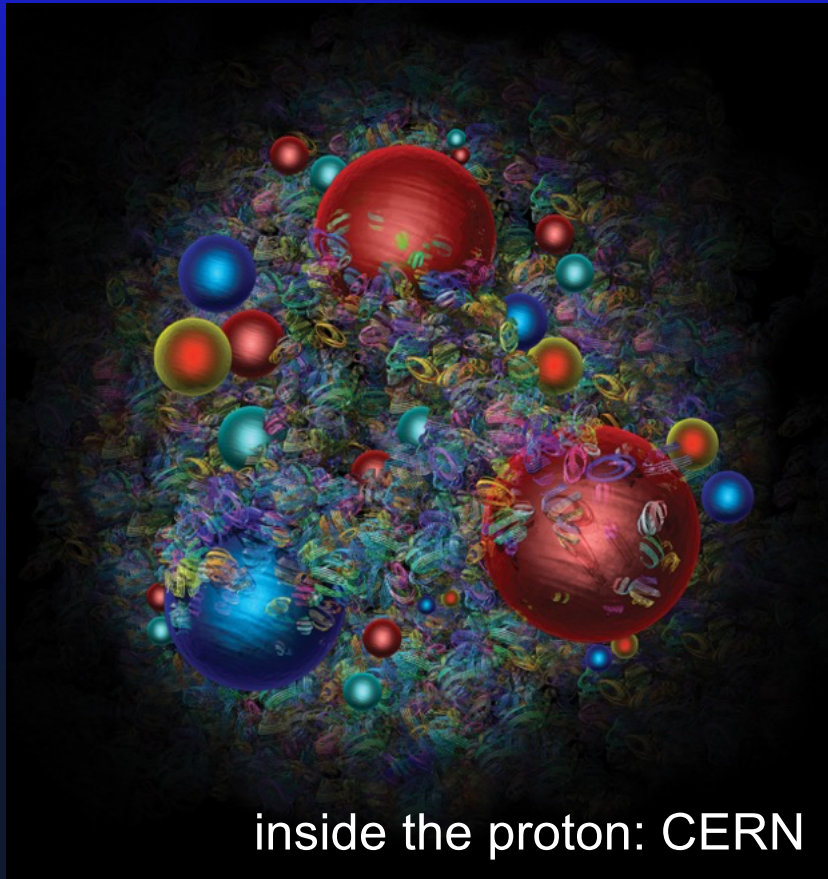
Gordon Baym
University of Illinois
Urbana, Illinois
&
iTHEMS RIKEN



東京大学
May 29, 2024



The Electron-Ion Collider: the ultimate electron microscope

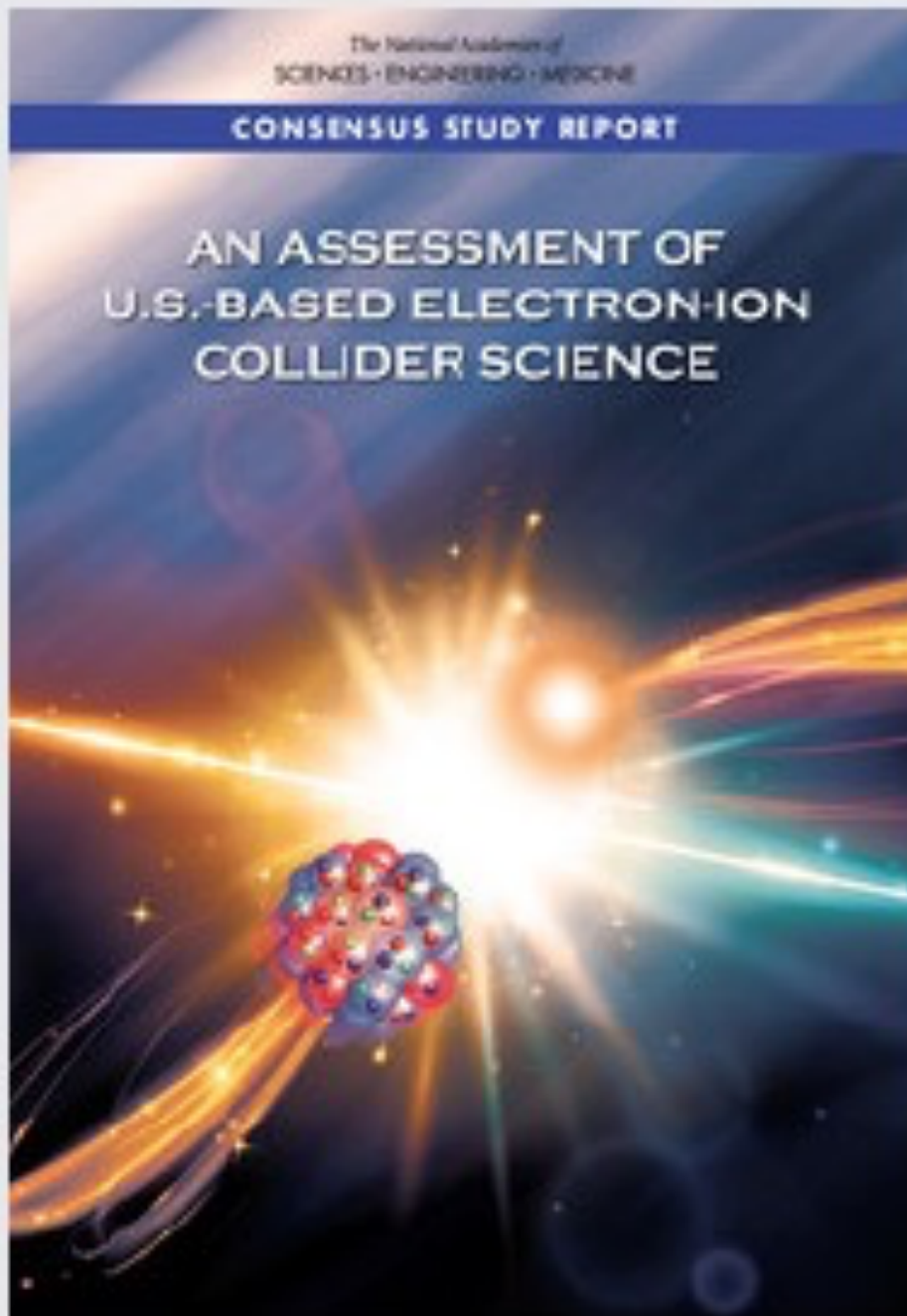


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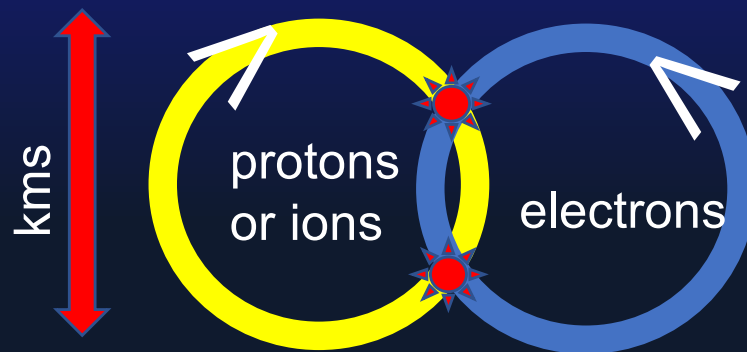
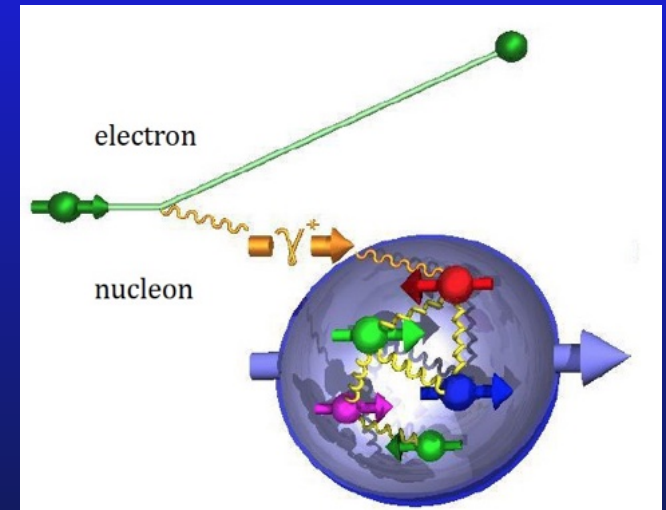
National Academy of Sciences
– National Research Council
report, July 2018, on the
science of an EIC

[Download pdf from
https://www.nap.edu/catalog/25171](https://www.nap.edu/catalog/25171)

The Electron-Ion Collider

A very big accelerator -- colliding beams of electrons with beams of protons or heavier ions (atomic nuclei).

A giant electron microscope for peering at the quarks and gluons deep inside the nucleon and atomic nuclei. **QCD machine.**



Electron-ion center of mass energy:

$$\sqrt{s} \sim 28 \sim 140 \text{ GeV.}$$

High luminosity (event rate) and spin polarized beams!

Proton mass $\sim 1 \text{ GeV}$

*Electron microscope
Invented 1931*



ca. 1940

Nuclear physics far from being a solved problem

Atomic nuclei: building blocks of the everyday world:

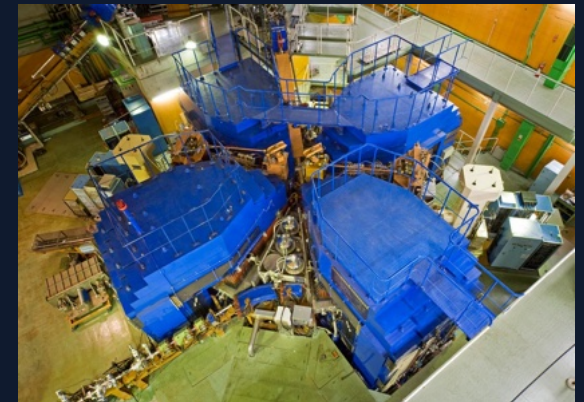


IUPAC Periodic Table of the Elements

IUPAC Periodic Table of the Elements																																															
1 H hydrogen (1.007, 1.008)																	18 He helium 4.003																														
3 Li lithium (6.938, 6.997)	4 Be beryllium 9.012	Key: atomic number Symbol name standard atomic weight											13 B boron (10.80, 10.83)	14 C carbon (12.00, 12.02)	15 N nitrogen (14.00, 14.01)	16 O oxygen (15.99, 16.00)	17 F fluorine 19.00	10 Ne neon 20.18																													
11 Na sodium 22.99	12 Mg magnesium (24.30, 24.31)											13 Al aluminium 26.98	14 Si silicon (28.08, 28.09)	15 P phosphorus 30.97	16 S sulfur (32.05, 32.08)	17 Cl chlorine (35.44, 35.46)	18 Ar argon 39.95																														
19 K potassium 39.10	20 Ca calcium 40.08	21 Sc scandium 44.96	22 Ti titanium 47.87	23 V vanadium 50.94	24 Cr chromium 52.00	25 Mn manganese 54.94	26 Fe iron 55.85	27 Co cobalt 58.93	28 Ni nickel 58.69	29 Cu copper 63.55	30 Zn zinc (65.38(2))	31 Ga gallium 69.72	32 Ge germanium 72.63	33 As arsenic 74.82	34 Se selenium (78.96(8))	35 Br bromine (79.90, 79.91)	36 Kr krypton 83.80																														
37 Rb rubidium 85.47	38 Sr strontium 87.62	39 Y yttrium 88.91	40 Zr zirconium 91.22	41 Nb niobium 92.91	42 Mo molybdenum 95.96(2)	43 Tc technetium	44 Ru ruthenium 101.1	45 Rh rhodium 102.9	46 Pd palladium 106.4	47 Ag silver 107.9	48 Cd cadmium 112.4	49 In indium 114.8	50 Sn tin 118.7	51 Sb antimony 121.8	52 Te tellurium 127.6	53 I iodine 126.9	54 Xe xenon 131.3																														
55 Cs caesium 132.9	56 Ba barium 137.3	57-71 lanthanoids	72 Hf hafnium 178.5	73 Ta tantalum 180.9	74 W tungsten 183.8	75 Re rhenium 186.2	76 Os osmium 190.2	77 Ir iridium 192.2	78 Pt platinum 195.1	79 Au gold 197.0	80 Hg mercury 200.6	81 Tl thallium (204.3, 204.4)	82 Pb lead 207.2	83 Bi bismuth 208.9	84 Po polonium	85 At astatine	86 Rn radon																														
87 Fr francium	88 Ra radium	89-103 actinoids	104 Rf rutherfordium	105 Db dubnium	106 Sg seaborgium	107 Bh bohrium	108 Hs hassium	109 Mt meitnerium	110 Ds darmstadtium	111 Rg roentgenium	112 Cn copernicium	113 Nh nihonium	114 Fl flerovium	115 Mc moscovium	116 Lv livermorium	117 Ts tennessine	118 Og oganesson																														
<table border="1"> <tr> <td>57 La lanthanum 138.9</td> <td>58 Ce cerium 140.1</td> <td>59 Pr praseodymium 140.9</td> <td>60 Nd neodymium 144.2</td> <td>61 Pm promethium</td> <td>62 Sm samarium 150.4</td> <td>63 Eu europium 152.0</td> <td>64 Gd gadolinium 157.3</td> <td>65 Tb terbium 158.9</td> <td>66 Dy dysprosium 162.5</td> <td>67 Ho holmium 164.9</td> <td>68 Er erbium 167.3</td> <td>69 Tm thulium 168.9</td> <td>70 Yb ytterbium 173.1</td> <td>71 Lu lutetium 175.0</td> </tr> <tr> <td>89 Ac actinium</td> <td>90 Th thorium 232.0</td> <td>91 Pa protactinium 231.0</td> <td>92 U uranium 238.0</td> <td>93 Np neptunium</td> <td>94 Pu plutonium</td> <td>95 Am americium</td> <td>96 Cm curium</td> <td>97 Bk berkelium</td> <td>98 Cf californium</td> <td>99 Es einsteinium</td> <td>100 Fm fermium</td> <td>101 Md mendelevium</td> <td>102 No nobelium</td> <td>103 Lr lawrencium</td> </tr> </table>																		57 La lanthanum 138.9	58 Ce cerium 140.1	59 Pr praseodymium 140.9	60 Nd neodymium 144.2	61 Pm promethium	62 Sm samarium 150.4	63 Eu europium 152.0	64 Gd gadolinium 157.3	65 Tb terbium 158.9	66 Dy dysprosium 162.5	67 Ho holmium 164.9	68 Er erbium 167.3	69 Tm thulium 168.9	70 Yb ytterbium 173.1	71 Lu lutetium 175.0	89 Ac actinium	90 Th thorium 232.0	91 Pa protactinium 231.0	92 U uranium 238.0	93 Np neptunium	94 Pu plutonium	95 Am americium	96 Cm curium	97 Bk berkelium	98 Cf californium	99 Es einsteinium	100 Fm fermium	101 Md mendelevium	102 No nobelium	103 Lr lawrencium
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The periodic table of chemical elements is over 150 years old. Are there further elements out there? Any of them stable?

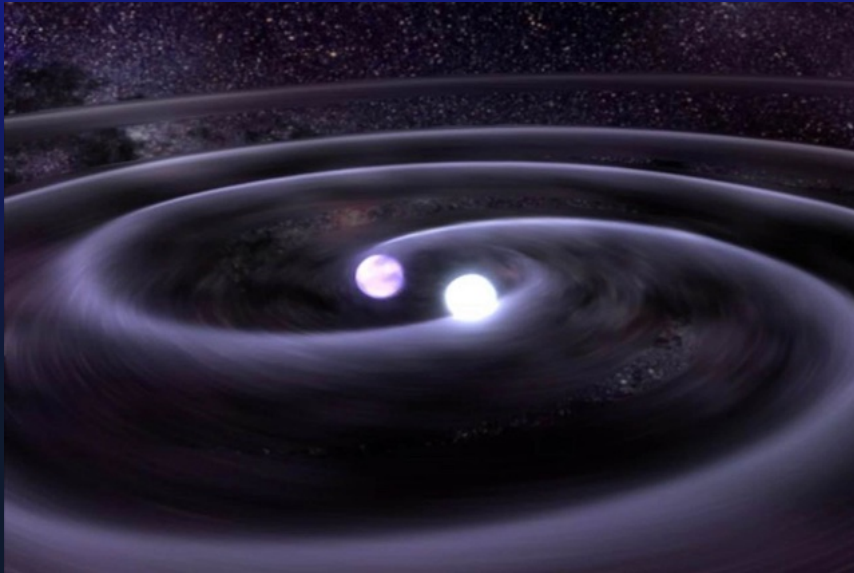
How many isotopes does each element have?
Answers from rare isotope accelerators (e.g., RIBF at RIKEN and FRIB at MSU) studying nuclei far from stability.



F = facility, RIB = rare isotope beam

RIBF

How to make the heavier elements? Answers, remarkably, from multimessenger studies of binary neutron star mergers: **Merger GW170817 observed on 17 Aug. 2017** by LIGO and Virgo (gravitational radiation), FERMI (gamma ray telescope) + some 70 other electromagnetic observatories.



Two neutron stars merging,
emitting gravitational radiation
and, post-merger, forming:

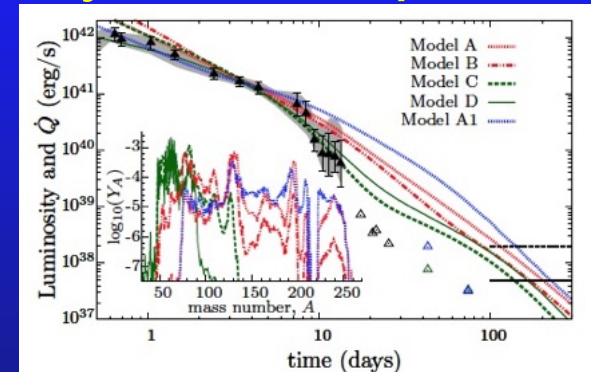


**Kilonova: neutron-rich
site of r-process**

Binary neutron star mergers likely site of heavy element production

r-process in *kilonova* =>
 earth-scale masses of
 Au, Ag, Pt, U... and Sr

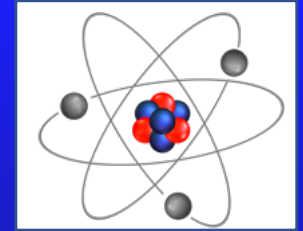
Kilonova
 N= 50, 82, 126



Atoms are made of electrons and nuclei.

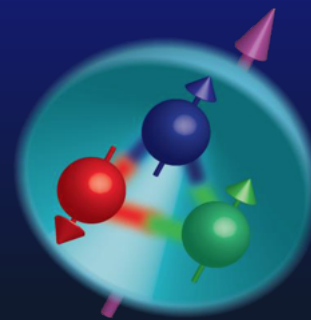
Similarly neutrons, protons, and nuclei are made of

quarks and gluons. But how? How to explain nucleon masses, spin, magnetic moments, etc. in terms of quarks and gluons?



Quarks = fractionally charged spin-1/2 fermions, baryon no. = 1/3, with internal SU(3) **color** degree of freedom.

Flavor	Charge/ e	Mass(MeV)
u	2/3	~2
d	-1/3	~5
s	-1/3	~ 94
c	2/3	~1280
b	-1/3	~4200
t	2/3	~175,000



$$\text{proton} = u + u + d$$

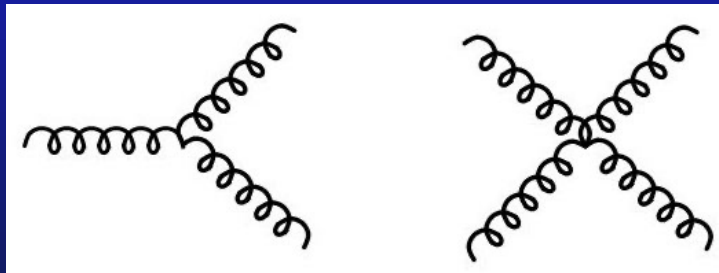
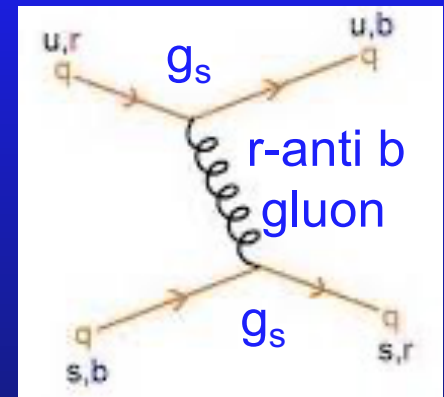
$$\text{neutron} = u + d + d$$

$$\pi^+ = u + \bar{d}, \text{ etc.}$$

Quark matter (baryonic) in the early universe at $t < 1$ microsec ($T > 100$ MeV), and in the deep interiors of heavier neutron stars.

Strong interactions – quantum chromodynamics

Quarks interact by exchanging gluons – massless vector bosons (like photon) with spin 1, and coming in 8 colors. **Gluons also interact with each other!!**



QCD

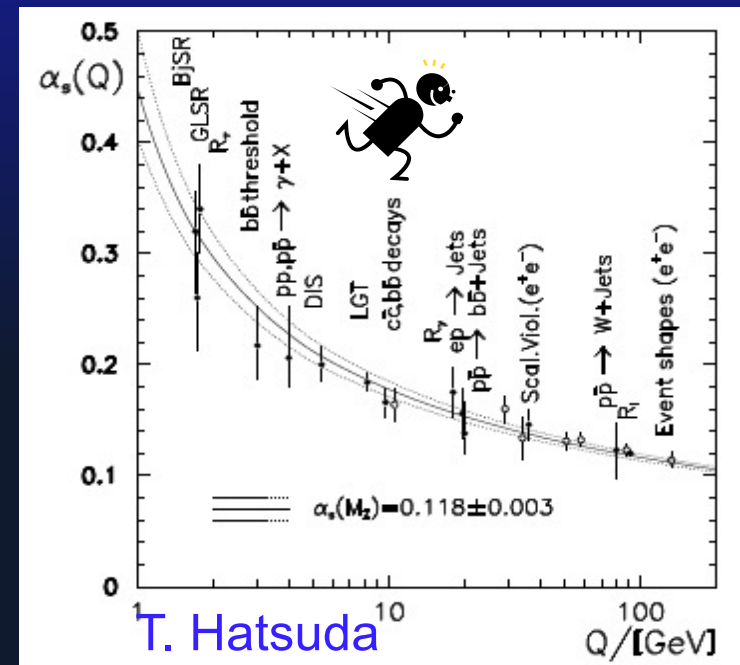
$$\alpha_s(\mu) = \frac{g_s^2}{4\pi} = \frac{6\pi}{(33 - 2N_f) \ln(\mu/\Lambda_{QCD})}$$

μ = energy scale $\Lambda_{QCD} \sim 340$ MeV

Asymptotic freedom as $\mu \rightarrow \infty$

(Even at Grand Unified (GUT) scale, 10^{15} GeV, g_s is not small: $\sim 1/2$; cf. electrodynamics:

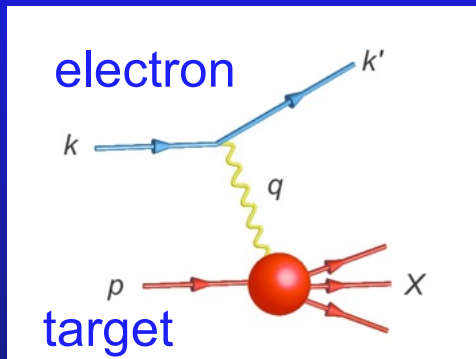
$$e^2/4\pi = 1/137 \Rightarrow e \sim 1/3)$$



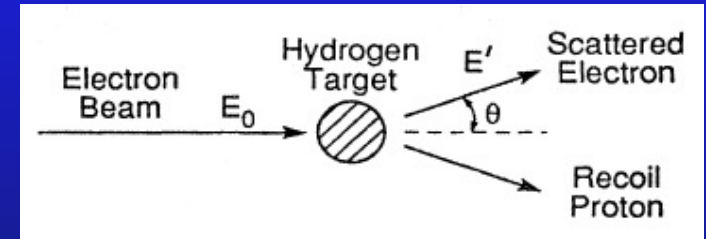
Running coupling constant

Electron scattering on nucleons and nuclei

Electron scattering on nucleons and nuclei: origins



Electrons (without internal structure) are precise probe of the complex structure of nucleons and nuclei.



First scattering of electrons on nuclei, Illinois Betatron 1951

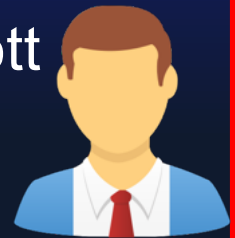
Al Hanson



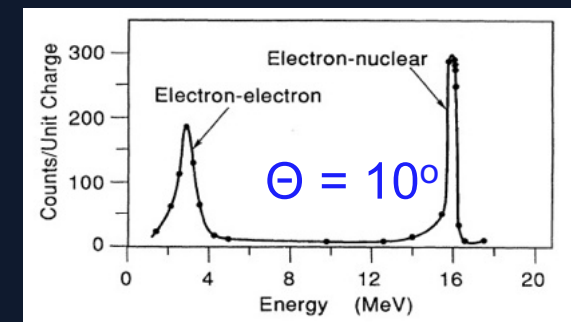
Ernie Lyman



and Merrill B. Scott



15.7 MeV electron beam on Be and Au foils





Donald Kerst, U of I, with the Betatron, ca. 1941

First scattering of electrons on protons and alpha particles

R. W. McAllister and Robert Hofstadter, 1956, with 187 MeV electrons at Stanford High Energy Physics Lab (HEPL) on H₂ and He gaseous targets, and then on polyacetylene (CH₂).



Found first hint of internal structure of the proton from the angular distribution:

proton radius $r_{\text{proton}} \sim 0.7 \text{ fm}$ (1fm = 10⁻¹³cm)

But is the inside of the proton a continuous “pudding” or are its constituents point particles?

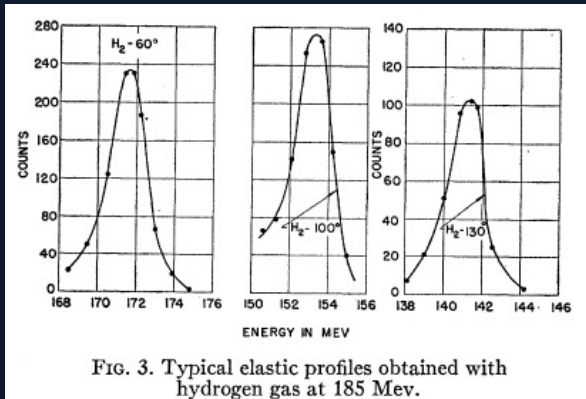
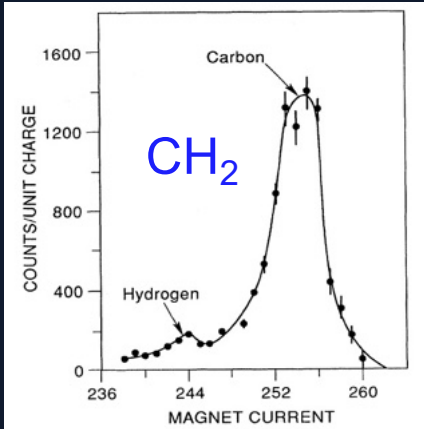
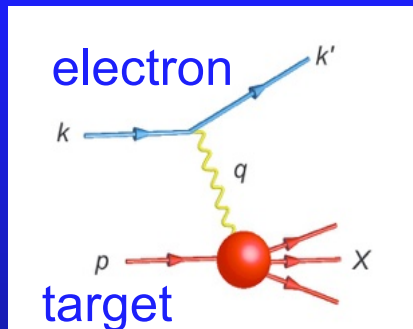


FIG. 3. Typical elastic profiles obtained with hydrogen gas at 185 Mev.

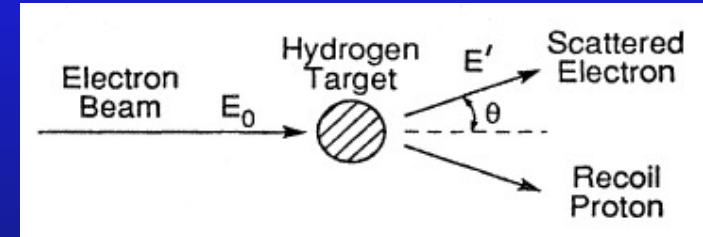
Elastic scattering on H₂ at various Θ





Kinematic variables in electron scattering from nucleons and nuclei

Electron scattering angle: θ



Electron energy transfer in lab: $\nu = E_e - E_e'$

3-momentum transfer from electron in lab: \mathbf{q}

4-momentum transfer squared: $Q^2 = \mathbf{q}^2 - \nu^2$
Larger $Q^2 =$ higher (transverse) resolution

Elastic scattering on target of mass m :

Energy conservation $(m^2 + q^2)^{1/2} = m + \nu \Rightarrow \nu = Q^2/2m$

Bjorken scaling variable: $x = Q^2/(2m_{\text{proton}} \nu)$
 x is like the “shutter speed”

The variables Q^2 and x define the landscape of electron scattering 15

Discovery of quarks 1967-73

Friedman, Kendall, and Taylor do **deep-inelastic scattering** (DIS) destroying proton target – at the Stanford Linear Accelerator Center (SLAC) 2 mile long, 20 GeV electron accelerator.

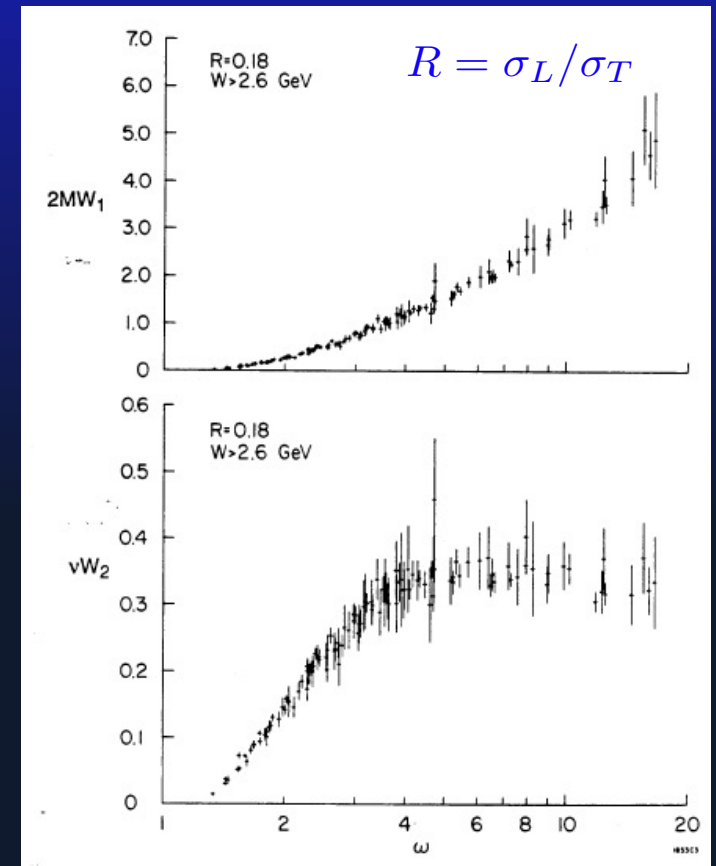
Measure electron angular cross section:

$$\sigma(E, E', \theta) = \frac{4e^4 E'^2}{Q^4} \left\{ W_2(\nu, Q^2) \cos^2 \frac{\theta}{2} + 2W_1(\nu, Q^2) \sin^2 \frac{\theta}{2} \right\}$$

For scattering from point particles inside proton, W_1 and νW_2 depend only on Bjorken scaling variable

$$x = \frac{Q^2}{2m_{proton}\nu}$$

Observe dependence on x only; shows that **proton is made of point particles**



Bjorken scaling indicating quark structure: $\omega = 1/x$ ($W =$ mass of recoiling target)

Proposal of quarks as mathematical model 1964

Volume 8, number 3

PHYSICS LETTERS

1 February 1964

A SCHEMATIC MODEL OF BARYONS AND MESONS *

M. GELL-MANN

California Institute of Technology, Pasadena, California

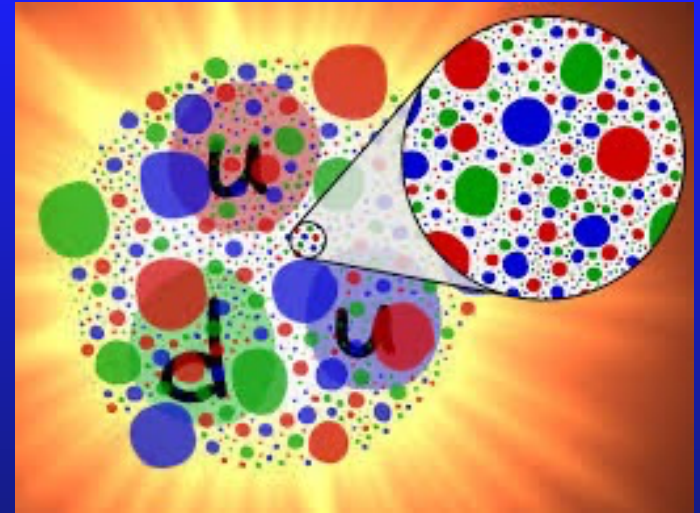
Ordinary

matter near the earth's surface would be contaminated by stable quarks as a result of high energy cosmic ray events throughout the earth's history, but the contamination is estimated to be so small that it would never have been detected. A search for stable quarks of charge $-\frac{1}{3}$ or $+\frac{2}{3}$ and/or stable di-quarks of charge $-\frac{2}{3}$ or $+\frac{1}{3}$ or $+\frac{4}{3}$ at the highest energy accelerators would help to reassure us of the non-existence of real quarks.

Parton model (Feynman, 1969)

building on J.D. Bjorken:

Understand electron scattering in terms of *partons* -- quasiparticles



Given parton carries momentum p (in beam direction) -- E'_e
 a fraction x of the total target proton momentum p_{proton} . **Same x**

$$\frac{p_{\text{parton}}}{p_{\text{proton}}} = x = \frac{Q^2}{2m_{\text{proton}}\nu}$$

energy conservation
 in scattering on parton $\nu = \frac{Q^2}{2m_{\text{parton}}}$

$$m_{\text{parton}} = x m_{\text{proton}}$$

Measure x in “infinite momentum” frame, i.e., with the proton moving at (nearly) the speed of light.

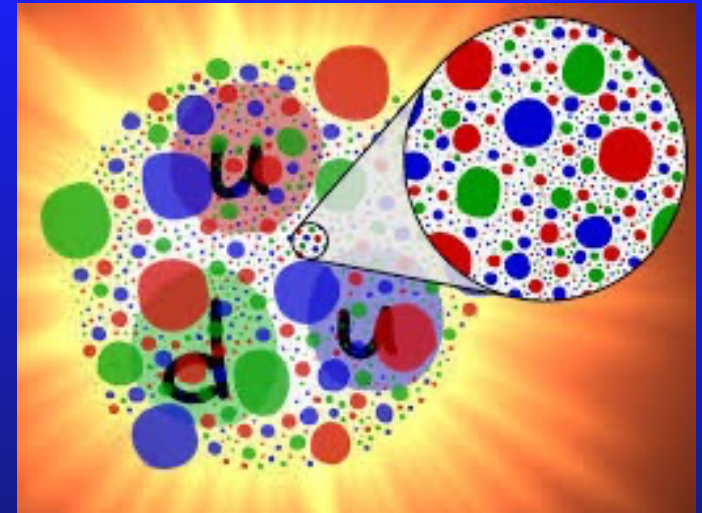
--Point partons identified with quarks, antiquark, and gluons -- all governed by QCD, with asymptotic freedom: **Bjorken and Paschos**

--Quarks are physical, not merely a tool: **Brodsky and Farrar**

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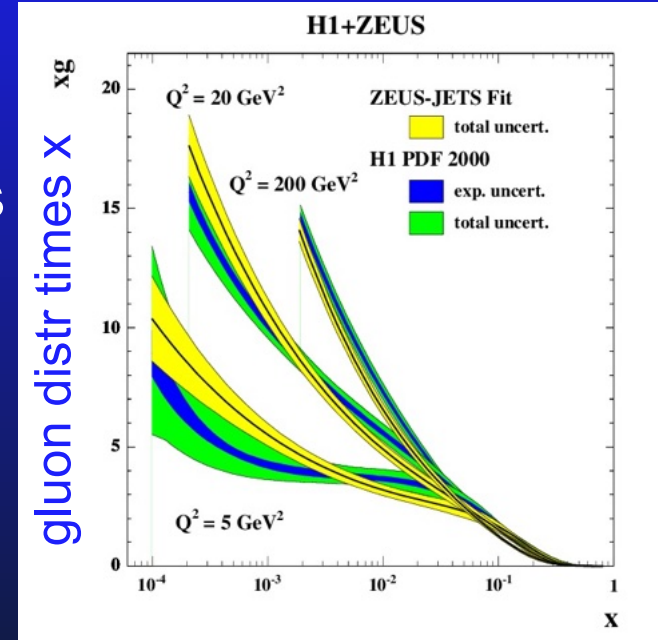
1991-2007 e scattering at HERA

(Hadron-Electron Ring Accelerator, Hamburg)

High-energy collisions of 27.5 GeV electron and positron beams (polarizable) with 920 GeV proton beams (unpolarized). No nuclear beams

HERMES (spin) fixed-target experiment.

HERA => great abundance of very low momentum ($x \ll 1$) gluons within nucleon.

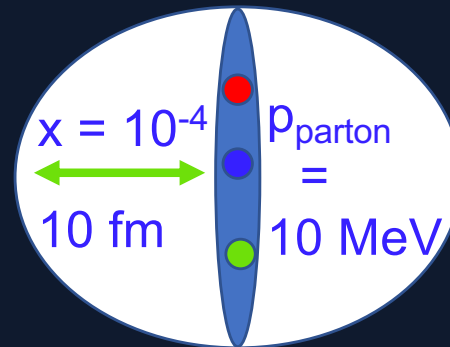


How can electrons probe electrically neutral gluons?

$g \rightarrow q + \bar{q}$ causing an effective electric dipole moment

View proton in frame in which proton is “slower” than ∞ momentum.

Heisenberg => low x “wee” partons stick out.



Lorentz contracted

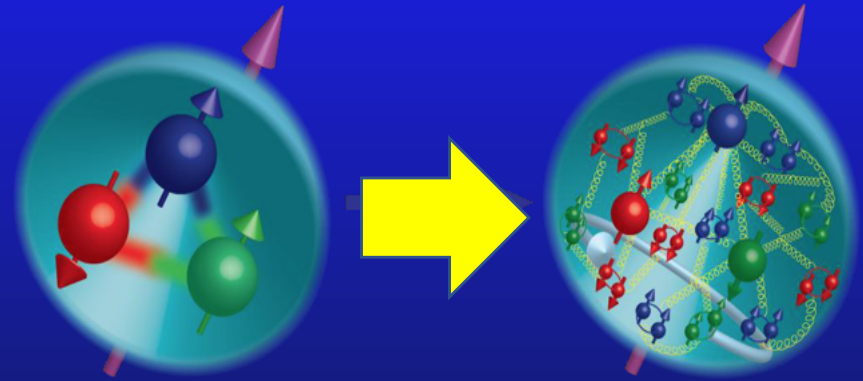


100 GeV proton

Physics with an electron-ion collider

Physics goals of an EIC

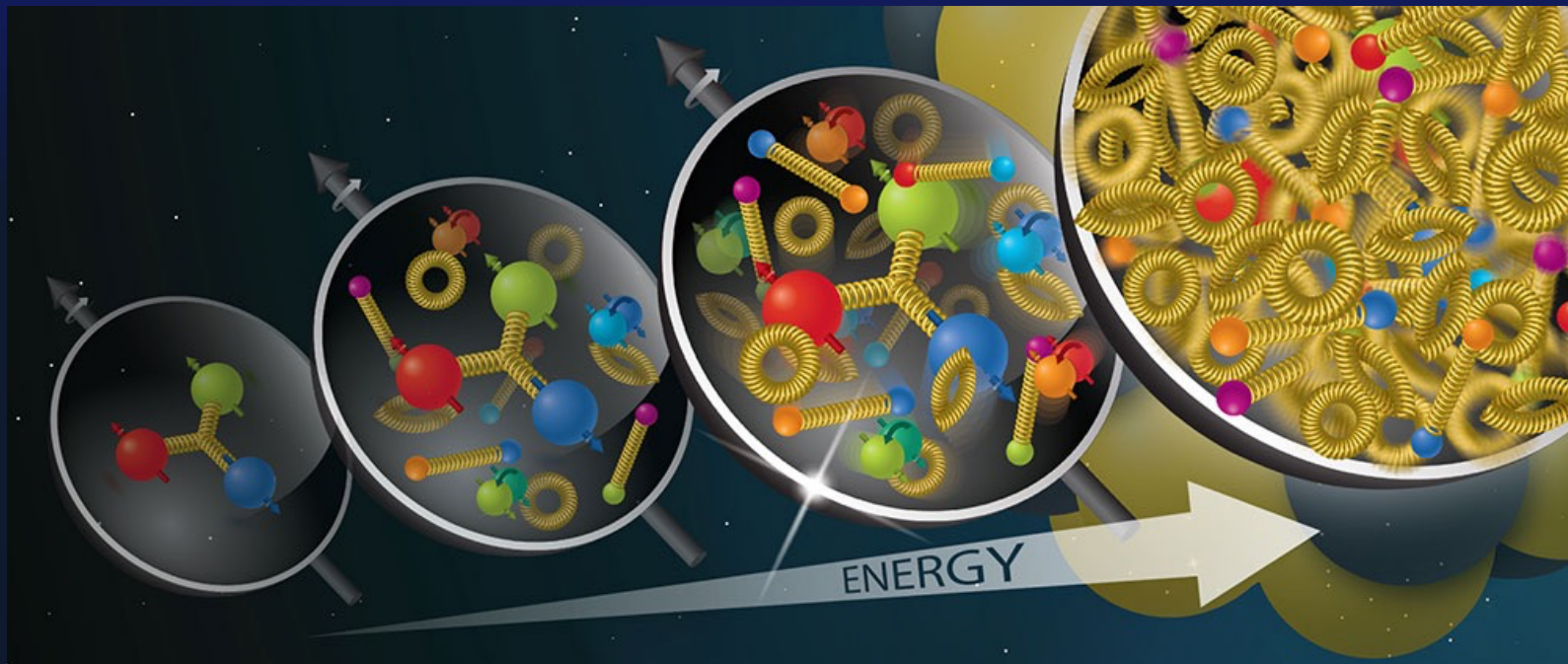
Our picture of the nucleon has evolved considerably from the first simple picture as a “bag” of three valence quarks. Have sea of quark-antiquark pairs (u, d, s) as well as cloud of gluons, all interacting!



1970s cartoon

Now

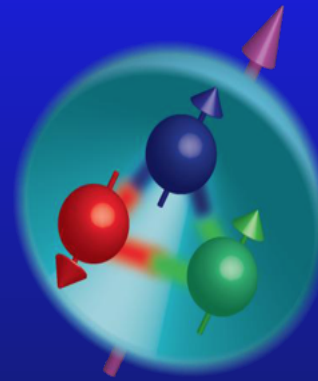
Z-E Meziani



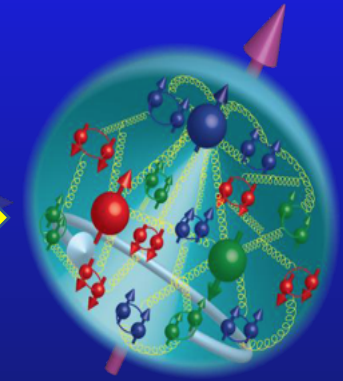
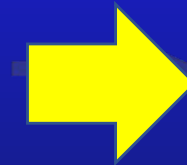
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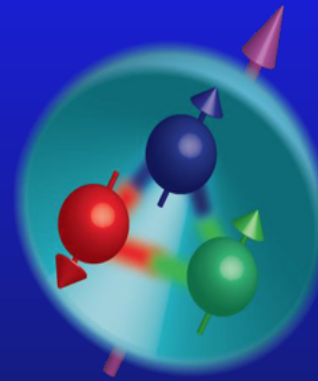
Z-E Meziani

Where are all the quarks and gluons -- in space and in momentum space? What is the many-body physics of all these interacting degrees of freedom.

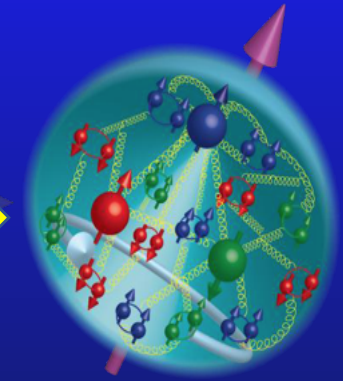
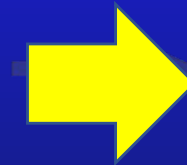
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Z-E Meziani

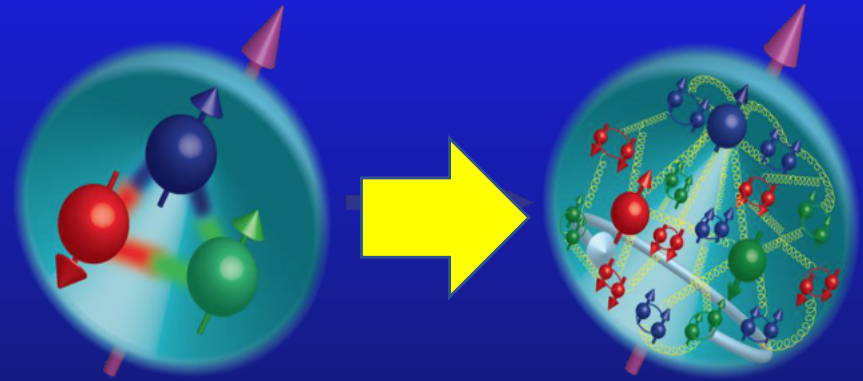
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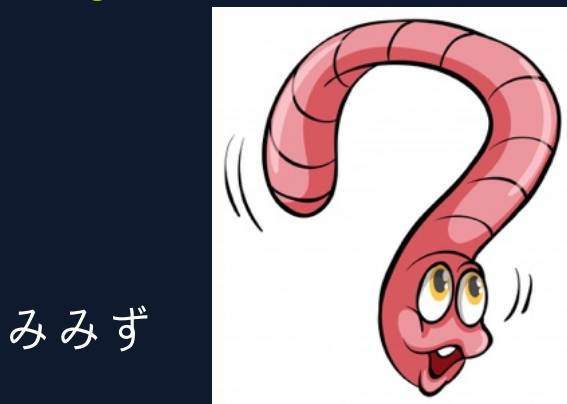


1970s cartoon

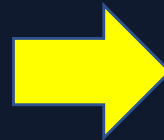
Now

Z-E Meziani

Where are all the quarks and gluons -- in space and in momentum space? What is the many-body physics of all these interacting degrees of freedom?



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FEI.com: Philippe Crassous

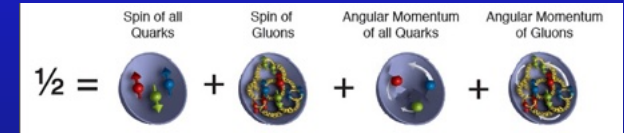
Sketch of worm vs. modern electron microscope picture (deep ocean worm)

Basic science questions for an EIC

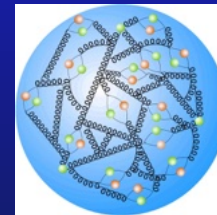
How does the nucleon get its mass?



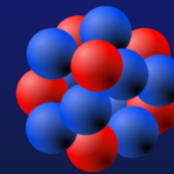
How does the spin of the nucleon arise from its elementary quark and gluon constituents?



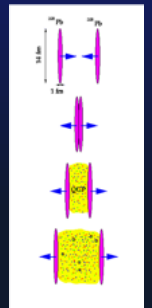
What are the emergent properties of dense systems of gluons?



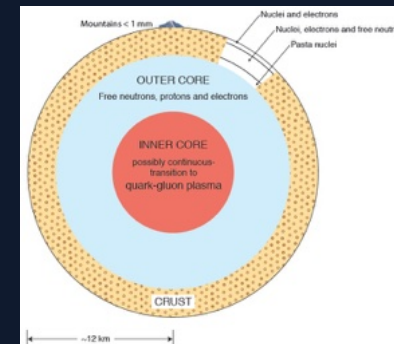
What is the internal structure of nuclei?
How do nuclei differ from being a simple collection of nucleons?



What is the “initial state” in ultrarelativistic heavy-ion collisions?



How does dense matter crossover from nucleonic degrees of freedom to quark degrees of freedom at higher density – application to neutron stars

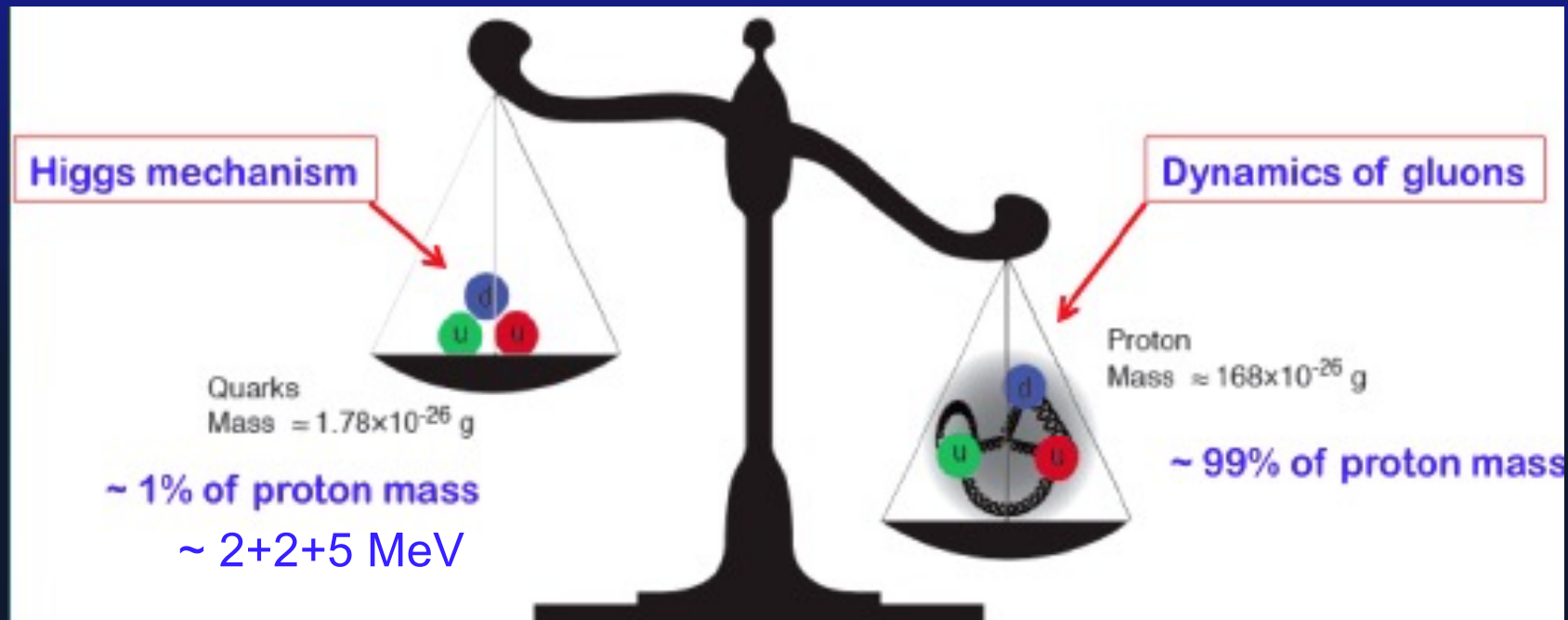


Signals of beyond-the-standard-model physics? Effects of new particles and forces?

How does a nucleon get mass?

$m_{\text{proton}} = 938 \text{ MeV}$ is almost 100 times greater than the masses of its valence quarks ($u+u+d \sim 2+2+5 \text{ MeV}$).

Cannot be understood in terms of the Higgs mechanism!!
Higgs => I should weigh $\sim 750 \text{ gm}$

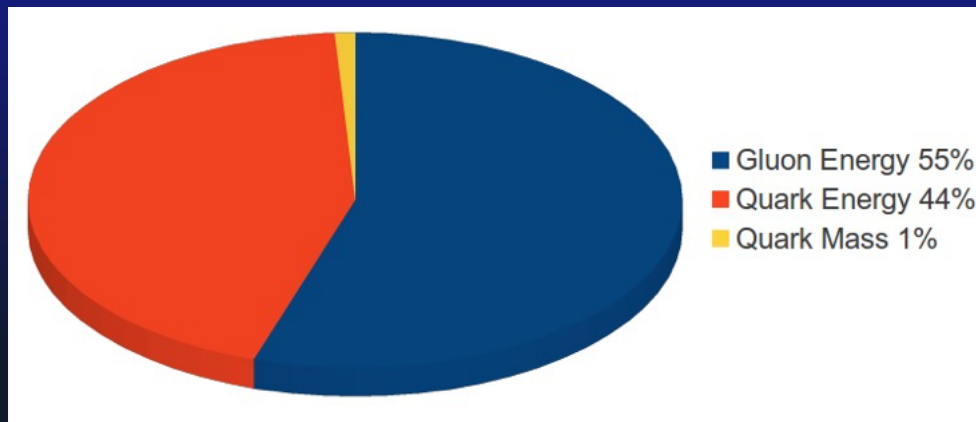


In atoms, mass = total mass of constituents minus binding energy (energy released in chemical reactions) < mass of constituents.
In nucleons, mass >> total mass of constituents.

Zero point energy => mass of the nucleon

Naively, the mass of the nucleon arises from all quark, anti-quark and gluon kinetic energies – from the uncertainty principle when localizing excitations within the nucleon (radius, r_p).

$$\Delta E \sim \hbar c / r_p \text{ per quark or gluon.}$$



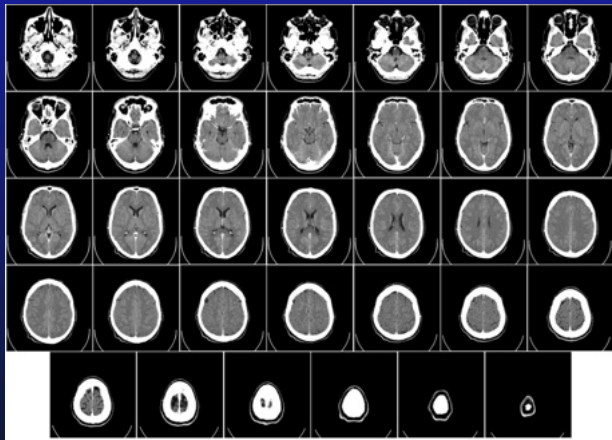
Energy distribution within nucleon

How are these components distributed (in space and momentum) in the nucleon?

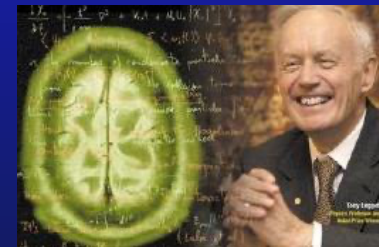
And how do nuclei differ from being a simple collection of nucleons – changes of quark and gluon distributions in nuclei?

Tomography

Determine internal structure of nucleon by measuring its momentum distributions (\vec{k}_\perp) transverse to the beam at varying x .
(Requires large luminosity.)

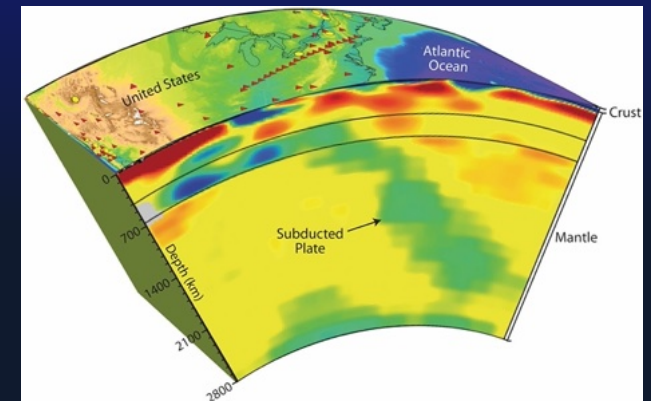


Computer assisted tomography of brain



A.J. Leggett

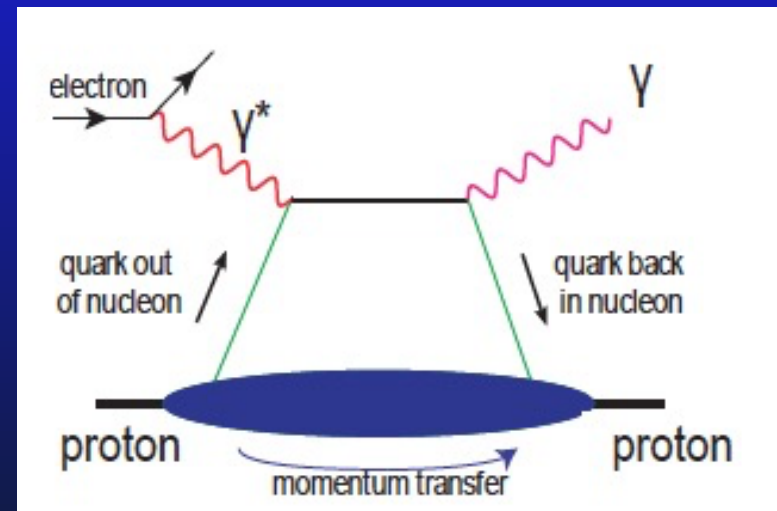
Seismic tomography, slices at varying depth; earthquakes as probes



Measure dependence of cross section on momentum transfer from electron to target => information about transverse position and momentum of struck quarks and gluons.

Measuring quark and gluon transverse momentum distributions

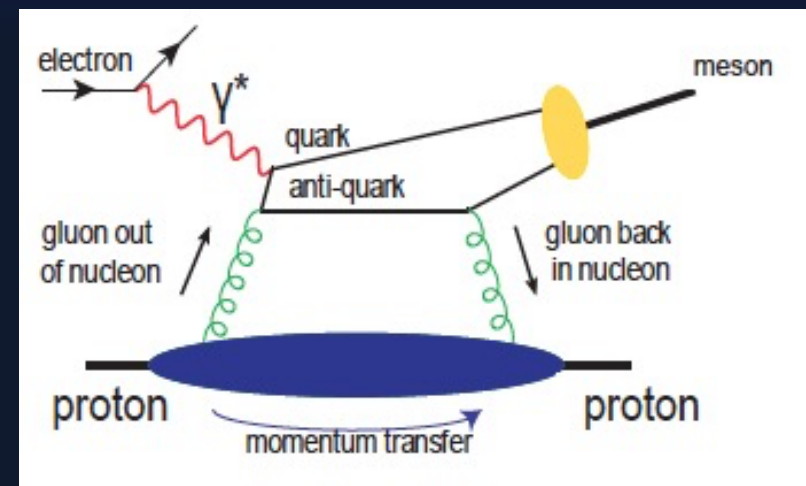
Real photon production γ :
Deeply virtual Compton scattering
sensitive to transverse position
of quarks.



Real meson production:
gives info on gluon distribution.

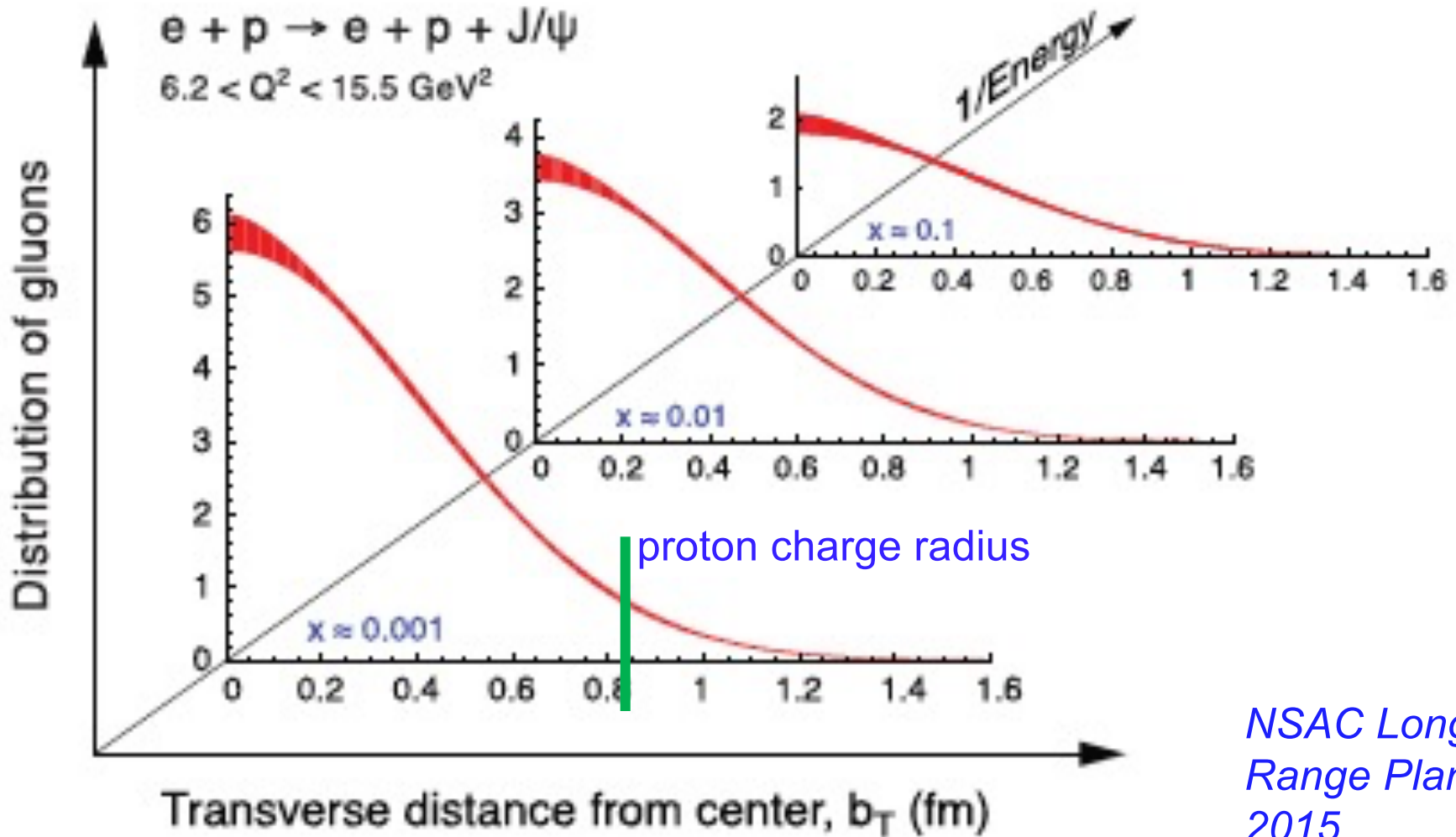
Ex., Heavy mesons J/ψ or Upsilon (Υ):

$$\gamma^* \rightarrow \bar{b}b \rightarrow \Upsilon$$



Expected transverse gluon distributions in space

Slices vs. transverse position b_T at various x

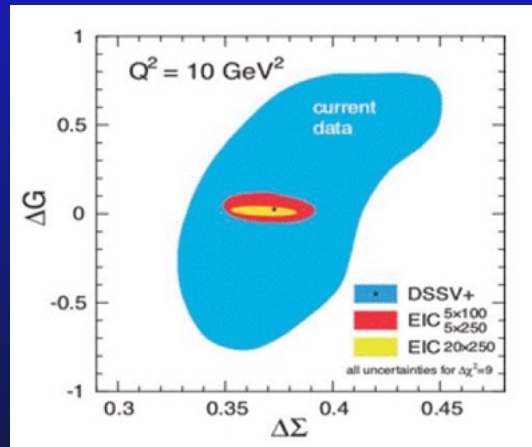


NSAC Long
Range Plan
2015

How do the quarks and gluons give the proton its spin?

Proton spin is the basis of nuclear magnetic resonance (MRI) imaging.

Spin crisis:



Current estimate of contributions to spin (in units of $\hbar/2$):

Quarks $\Delta\Sigma \sim 30\text{-}40\%$

Gluons $\Delta G \sim -70 \text{ to } +70\% \text{ ??}$
(RHIC pp)

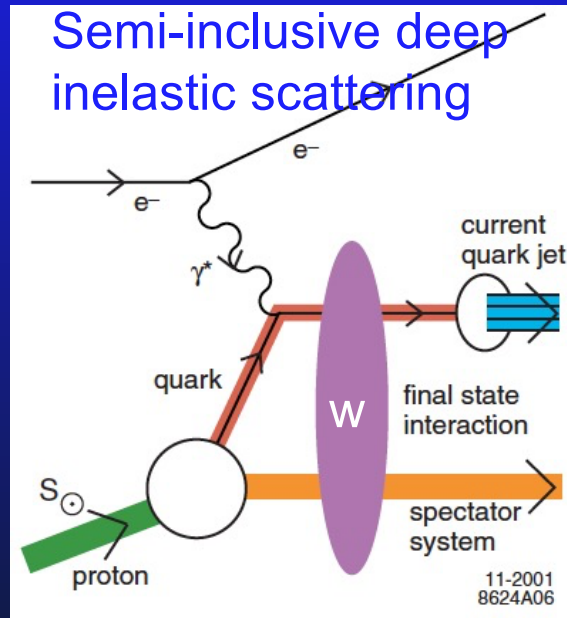
Orbital motion ?? (JLab - 12GeV)

Extract orbital contributions from transverse motion measured by spatial and momentum distributions of components in deep inelastic scattering -- **tomography**.

Extract gluon contributions from transfer of gluon polarization to $q\text{-bar } q$ pair, probed by polarized electrons.

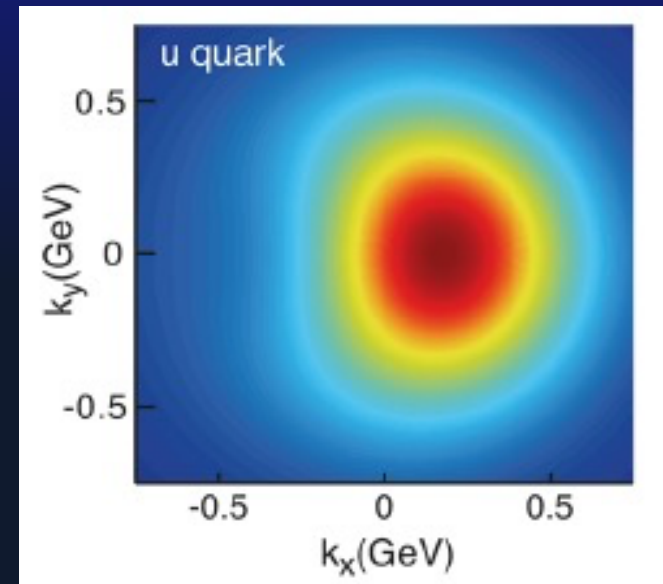
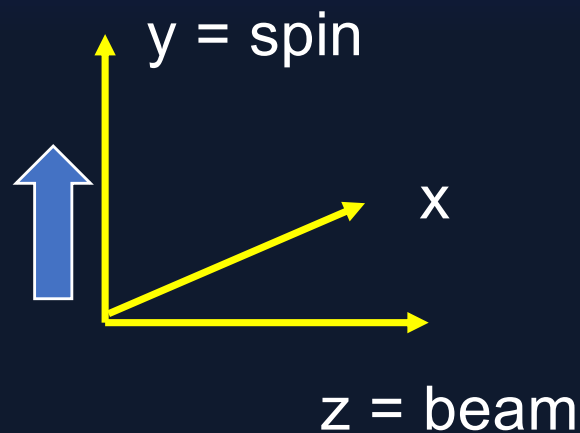
Polarized beams ($> 70\%$) in EIC critical !

Detecting quark orbital contributions to proton spin



Proton spinning (along y) with orbital motion as well leads to asymmetry of quark distributions in x direction.

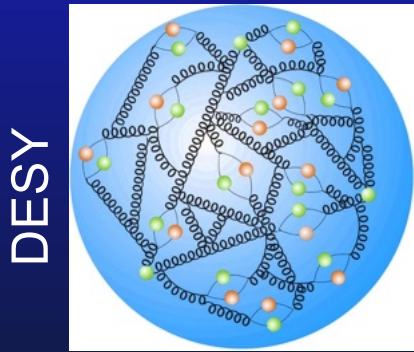
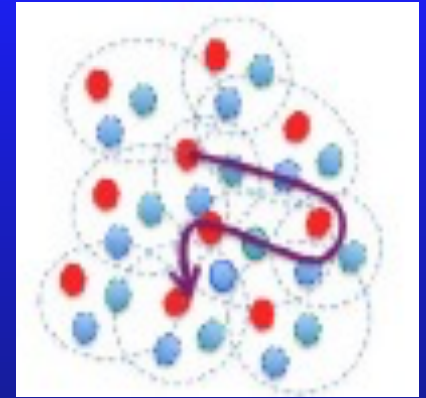
Beam along z .



Representative asymmetry of up-quark distribution for $x = 0.1$

Gluon physics

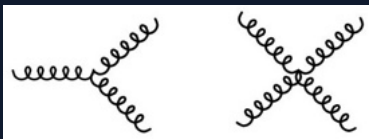
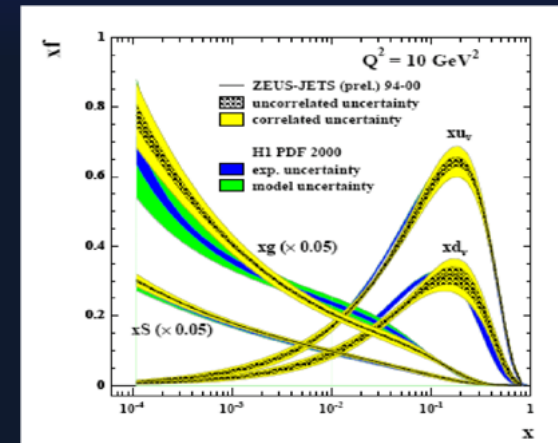
Gluons in nucleons and nuclei (as well as other hadrons) are like **dark matter** in the universe—unseen but crucial in holding matter together.



Nucleons and nuclei are in fact complex interacting many-body systems – not simply bags of free quarks and free gluons. Ex., **nuclei exhibit composite fermions—the nucleons**. Confinement!

“The most precise picture of the proton”

HERA => huge numbers of low momentum gluons in the nucleon -- at low x ($<10^{-4}$).
 Low momentum sector (“wee”) partons dominated by strongly interacting gluons!. The gluon field

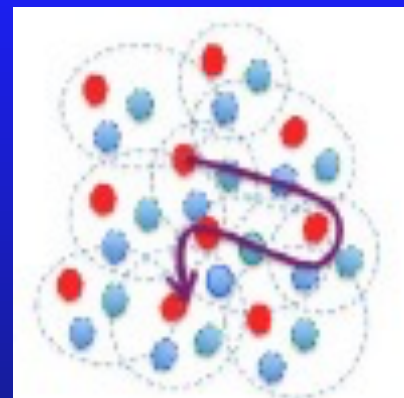


is highly non-linear!

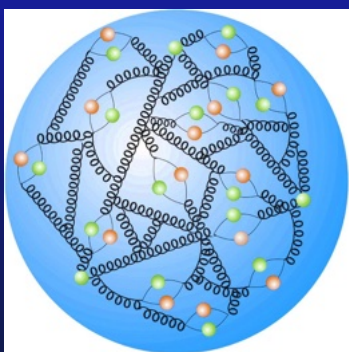
A new many-body system! New emergent phenomena?

Gluon physics

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DESY

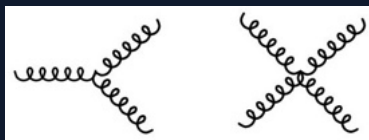


Nucleons and nuclei are in fact complex interacting many-body systems – not simply bags of free quarks and free gluons. Ex., **nuclei exhibit composite fermions—the nucleons**. Confinement!

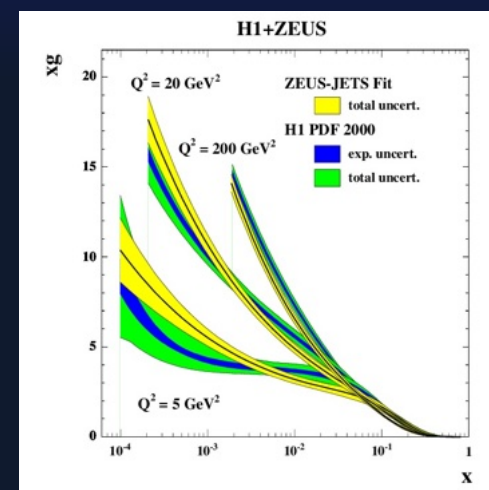
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A new many-body system! New emergent phenomena?

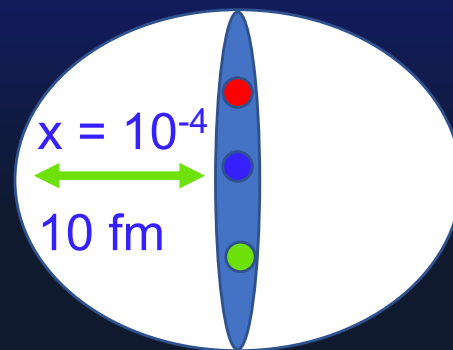
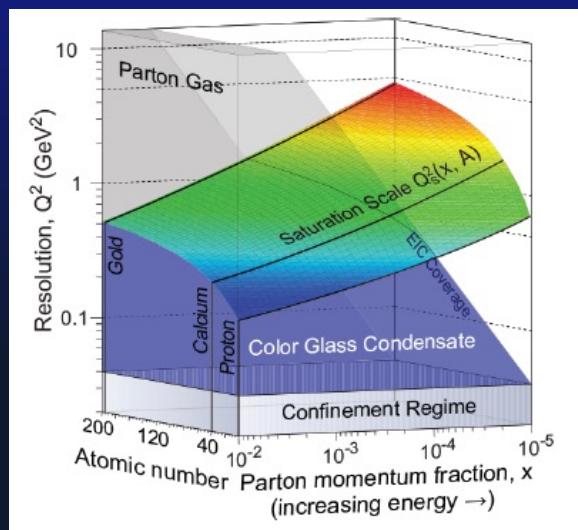


Gluon self-interactions become important at small x where there are many many gluons

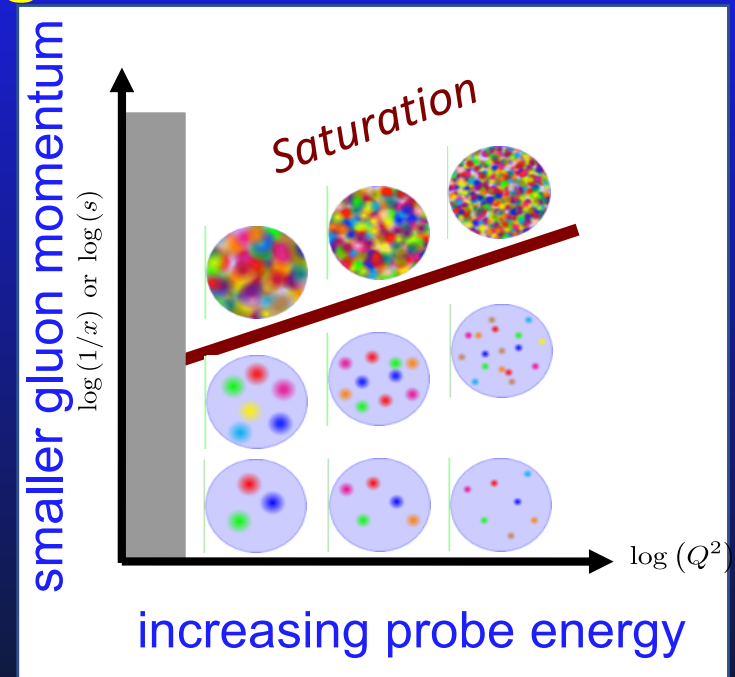
Scale of saturated gluonic matter: Q_s

At HERA (318 GeV c.m.) $Q_s^2 \sim 1 \text{ GeV}^2$

At EIC ~ 1 -few GeV^2 , growing with A .



Lorentz contracted
100 GeV proton



First approximation, dense cloud of gluons forms a **Bose condensate** –
“color glass condensate.”

Excitations of saturated gluonic matter? Topology?

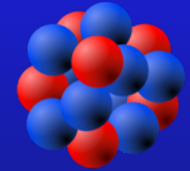
Evidence of non-linearity from STAR at RHIC: PRL 129, 092501 (2022)

Connections with nuclear physics elsewhere

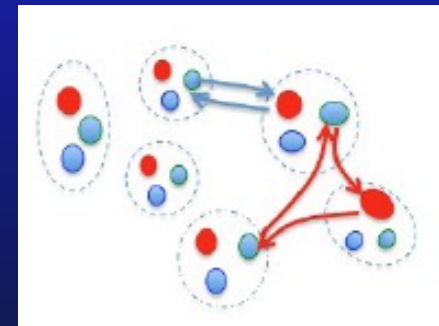
structure of the nucleon,
heavy ion
&
neutron star physics

What role do quarks play within laboratory nuclei?

How does the picture of individual non-overlapping nucleons within familiar nuclei break down?



Standard picture is in terms of nucleonic degrees of freedom interacting via forces. More microscopically have quark exchanges, leading to correlations among nucleons. How do nuclei differ from being a simple collection of nucleons?



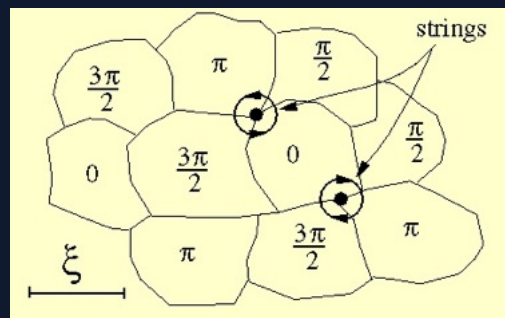
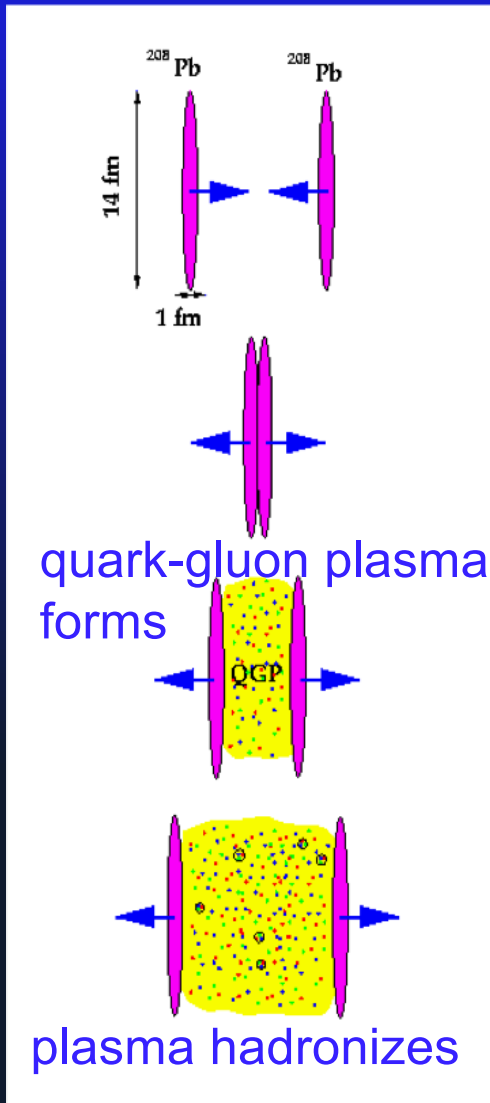
Where are all the quarks and gluons -- in space and in momentum space? What is the many-body physics of all these interacting degrees of freedom?

EIC should lead to new perspectives on the nuclear science being studied at other facilities, e.g., FRIB, and JPARC

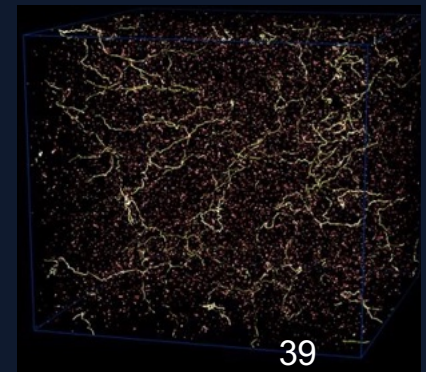
Connections to heavy ion collisions:

Saturated gluonic matter reachable at a sufficiently energetic EIC. Describes “initial state” in ultrarelativistic heavy ion collisions. **Bose-condensed gluonic matter** (color-glass condensate, ...). Condensate is metastable, decaying into quark-gluon plasma.

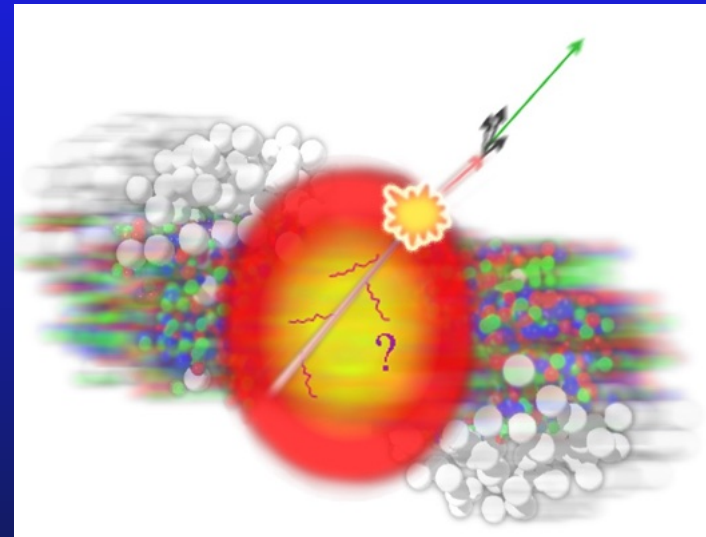
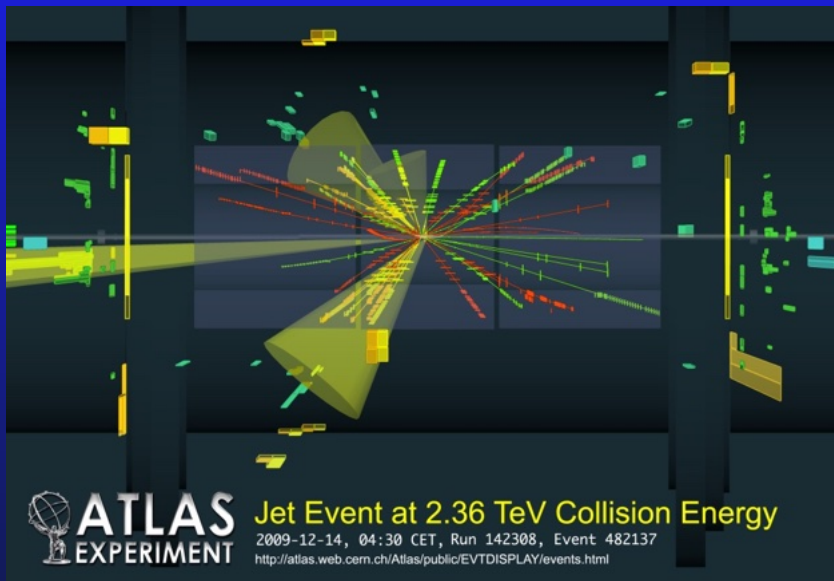
As in the early universe, expect topological defects, e.g., handedness asymmetry of produced $q \bar{q}$ pairs (chiral magnetic effect) related to the structure of the color field in saturated gluonic matter.



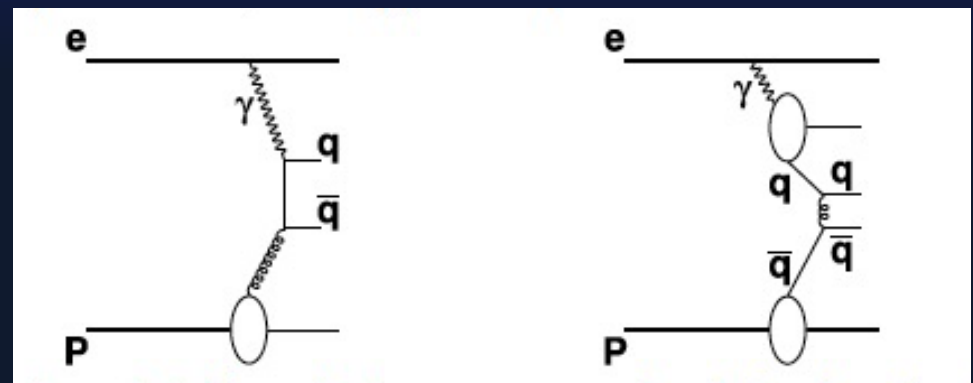
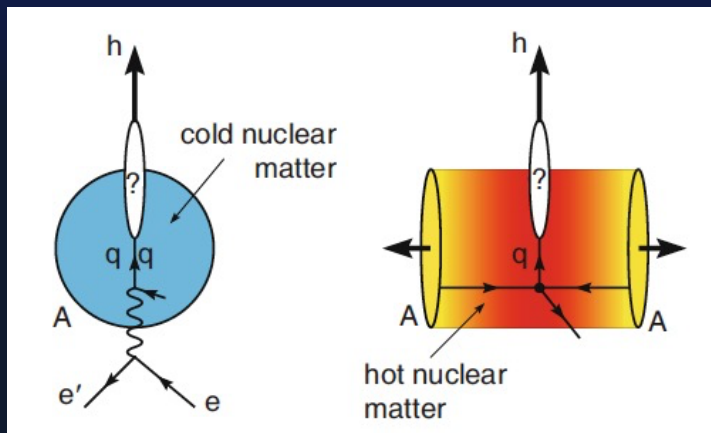
cosmic strings



Particle jets in electron-ion and ion-ion collisions



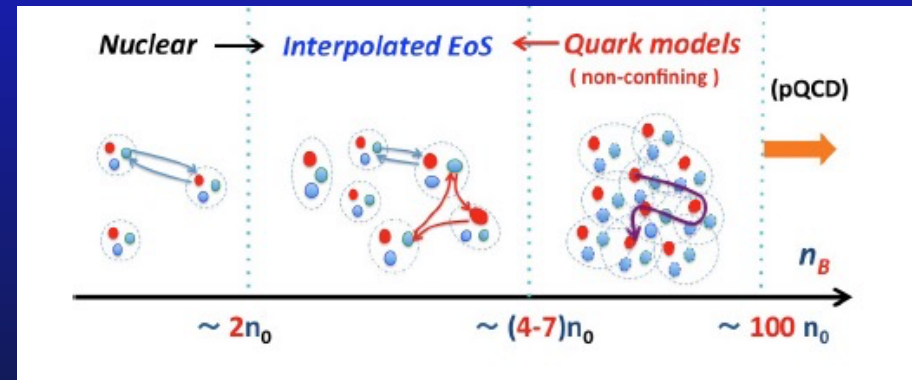
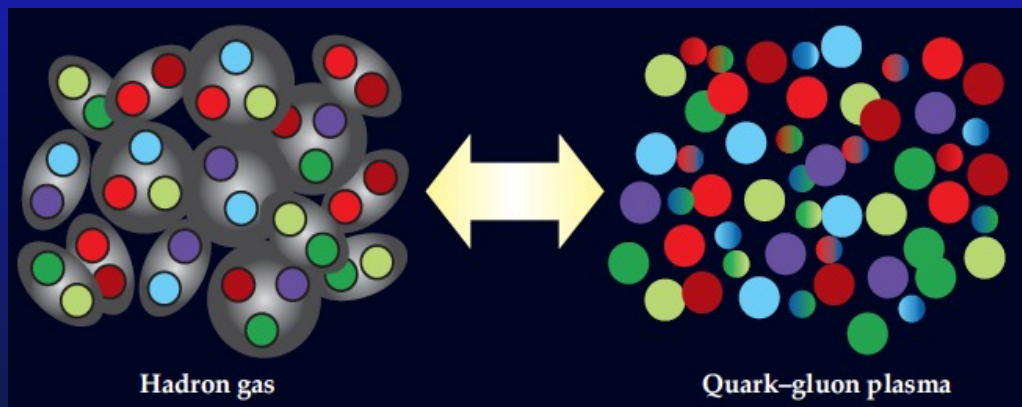
asymmetric jet in AA collision



Jets probe quark and gluon distributions

Dense matter and neutron stars:

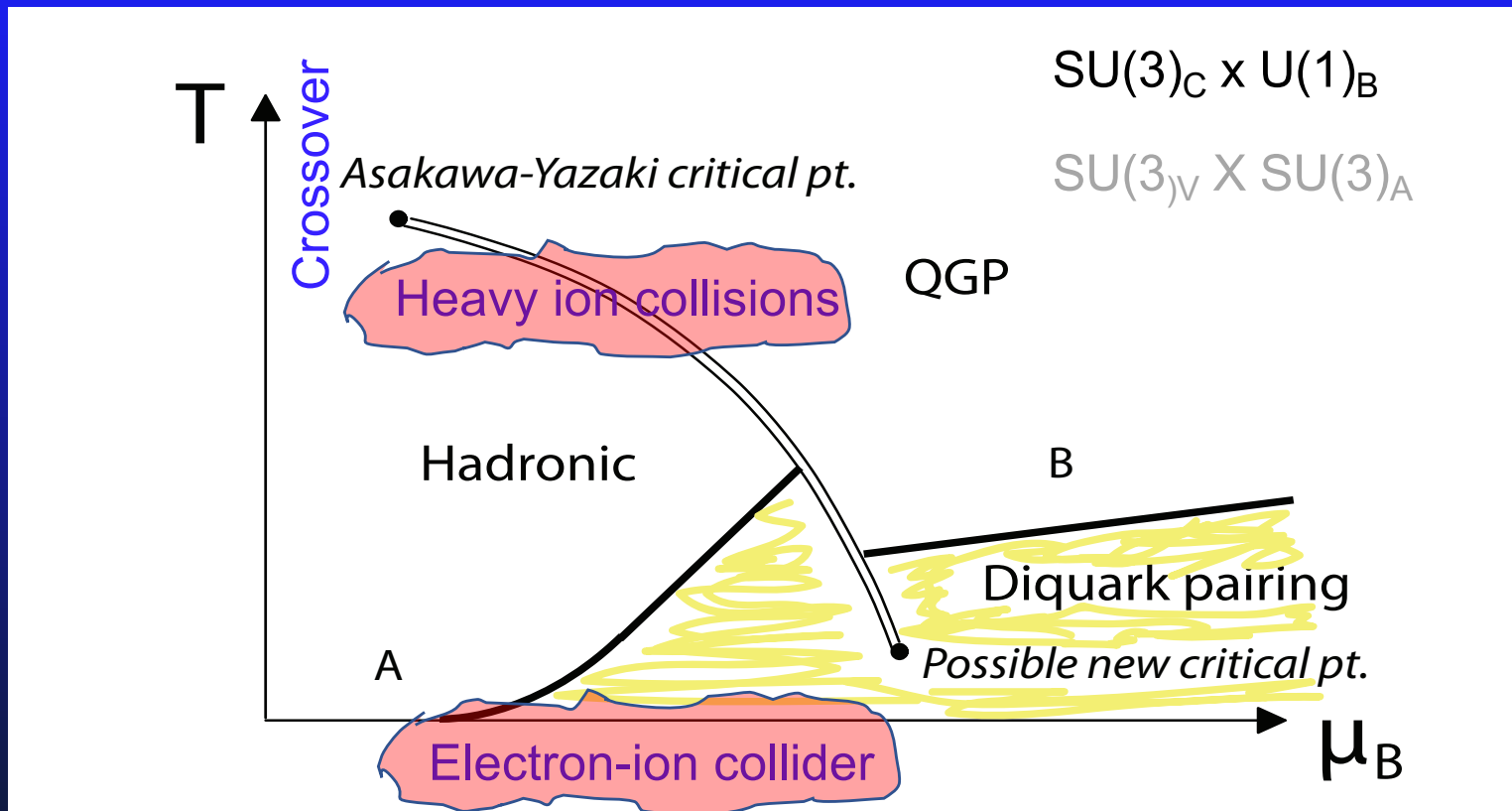
Study transition from cold nuclear matter to quark matter – vital for neutron stars. What is energy density vs. baryon density?



Expect “smooth” transition from nucleons to quarks

Gluon (and quark) distributions in nuclei at finer and finer scales should shed light on transition from nucleonic to quark degrees of freedom as density of matter increases.

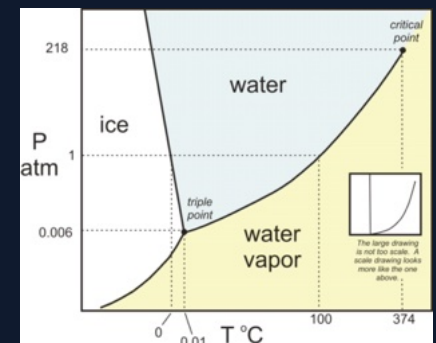
(Can mapping of energy-momentum tensor (stress-energy tensor) in eA collisions reveal pressure vs. baryon density in dense matter?)



Critical points similar to those in liquid-gas phase diagram (H_2O). **Neither critical point required!!**

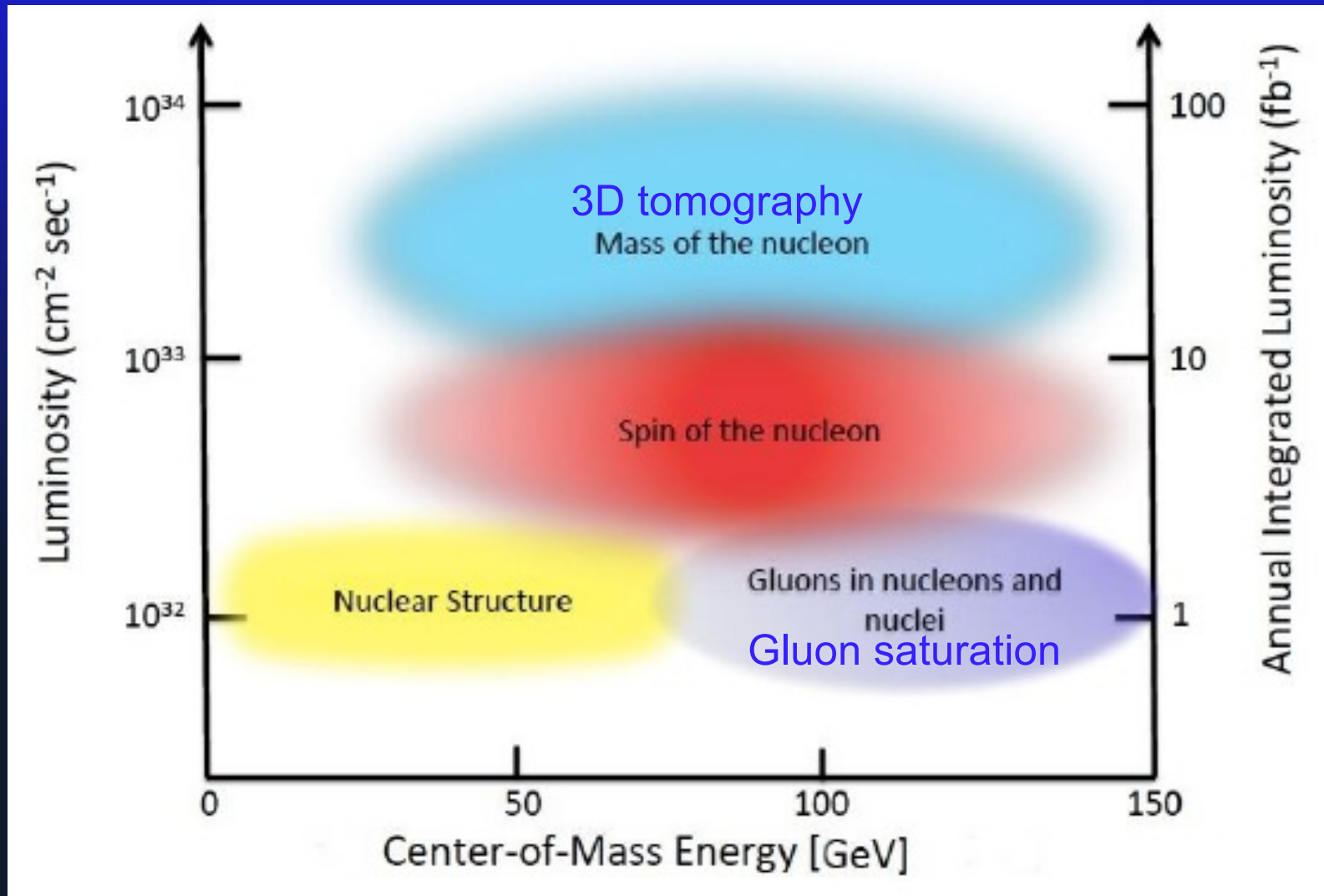
Can go continuously from A to B around the upper critical point. Liquid-gas phase transition.

In lower shaded region have BCS pairing of nucleons, of quarks, and possibly other states (meson condensates). Different symmetry structure than at higher T .



The accelerator

(electron-proton) c.m. energy - luminosity landscape



(Luminosity measures the rate of collisions: $L\sigma = \text{event rate}$) 43

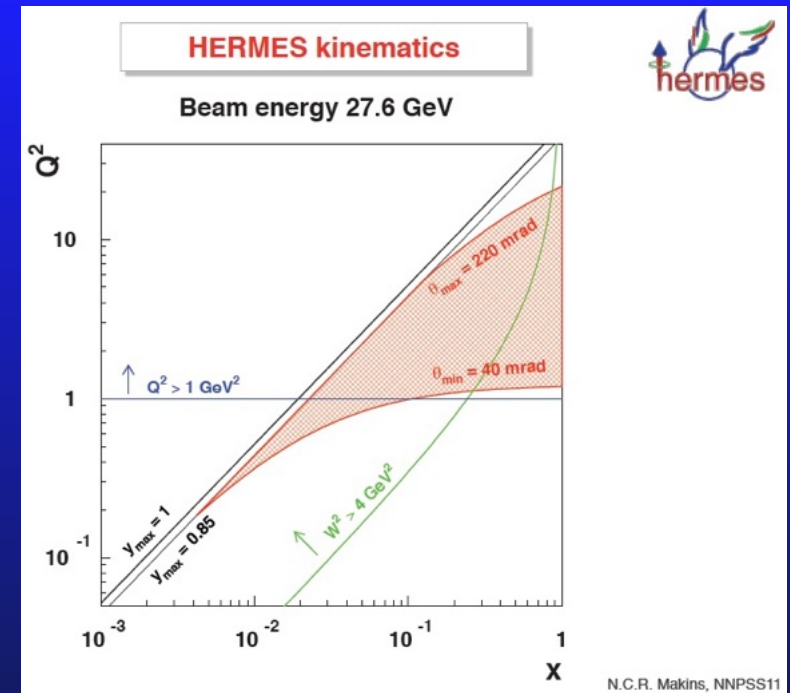
Kinematic range

$$Q^2 = x(1 - E_e'/E_e)s \leq x s$$

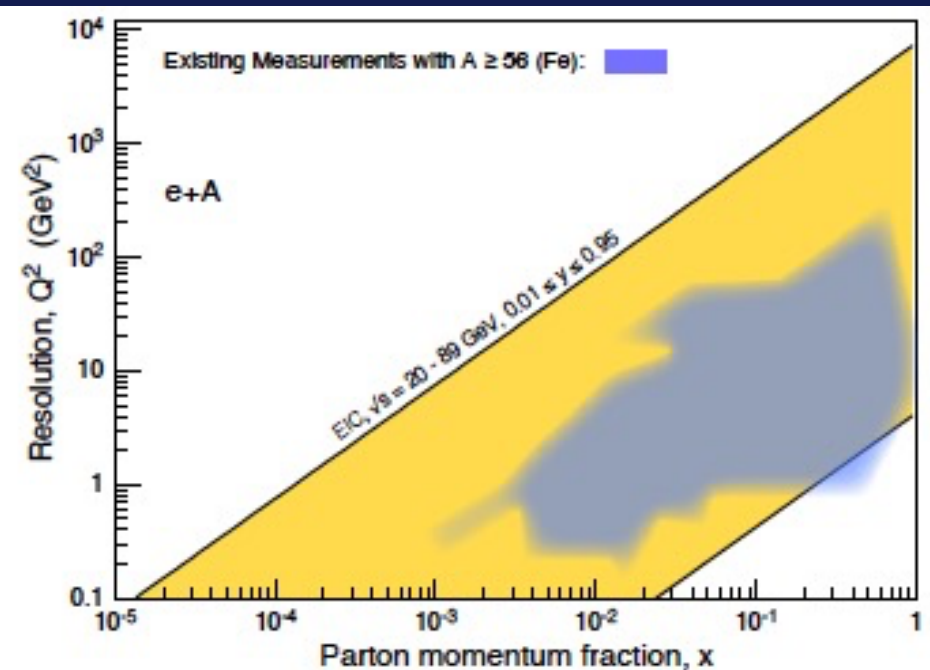
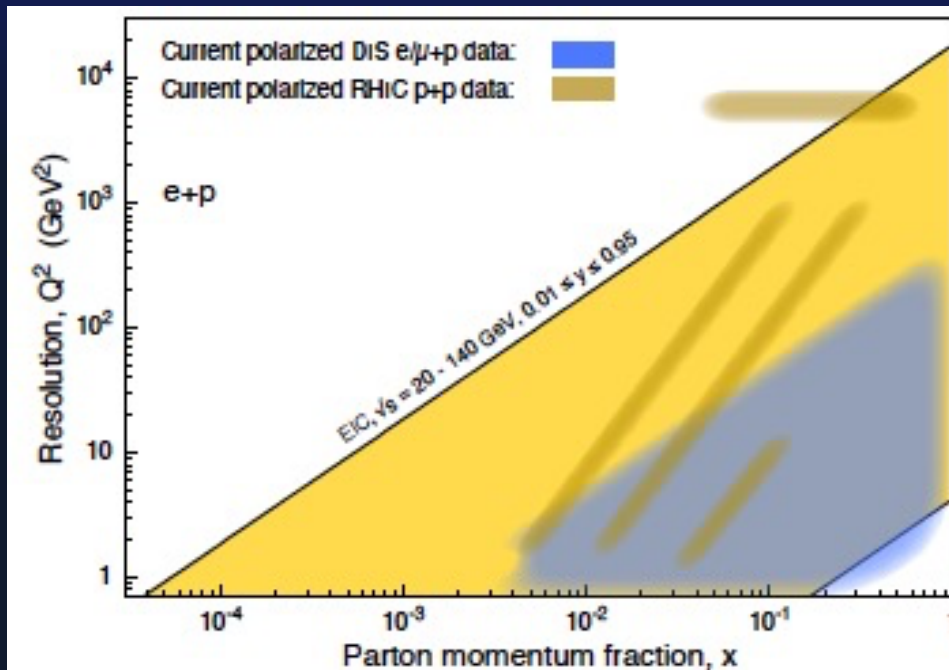
\sqrt{s} = center of mass collision energy

$x \gtrsim 10^{-4}$ for $Q^2 \sim 1 \text{ GeV}^2$, $\sqrt{s} \sim 100 \text{ GeV}$

($Q^2 \sim 1 \text{ GeV}^2$ lower limit for inelastic scattering)



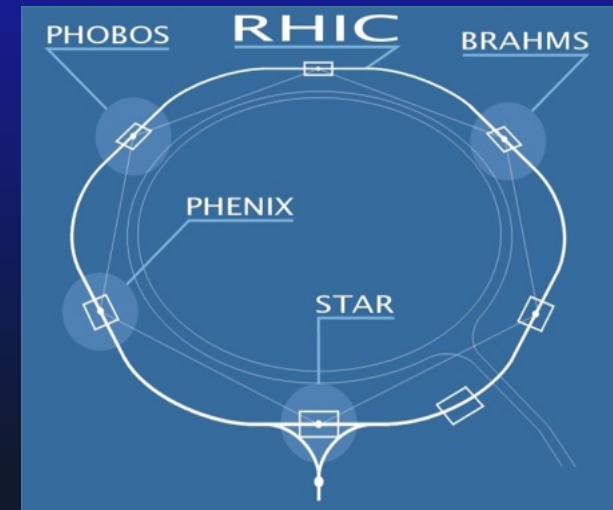
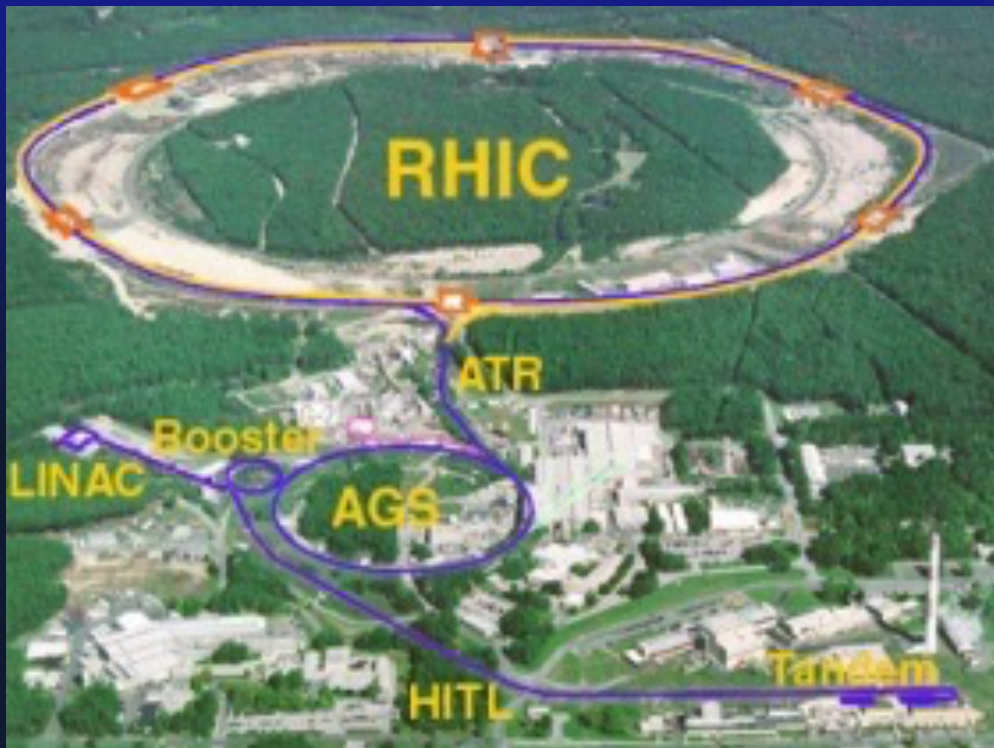
To reach small x region requires large beam energy



Relativistic Heavy Ion Collider (Brookhaven) since 2000 now becoming the EIC (joint project with Jefferson Lab)

RHIC: Colliding heavy ion beams 100 GeV/A

Colliding polarized proton beams $\sim 255 + 255$ GeV

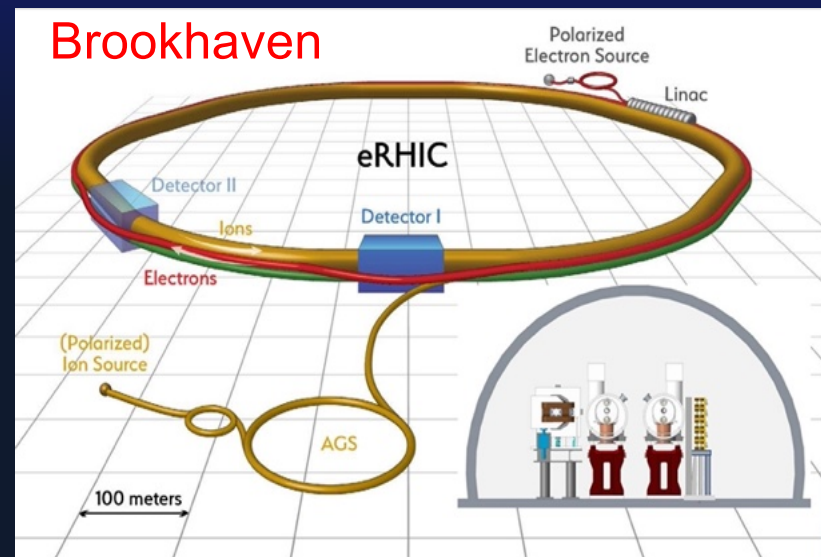
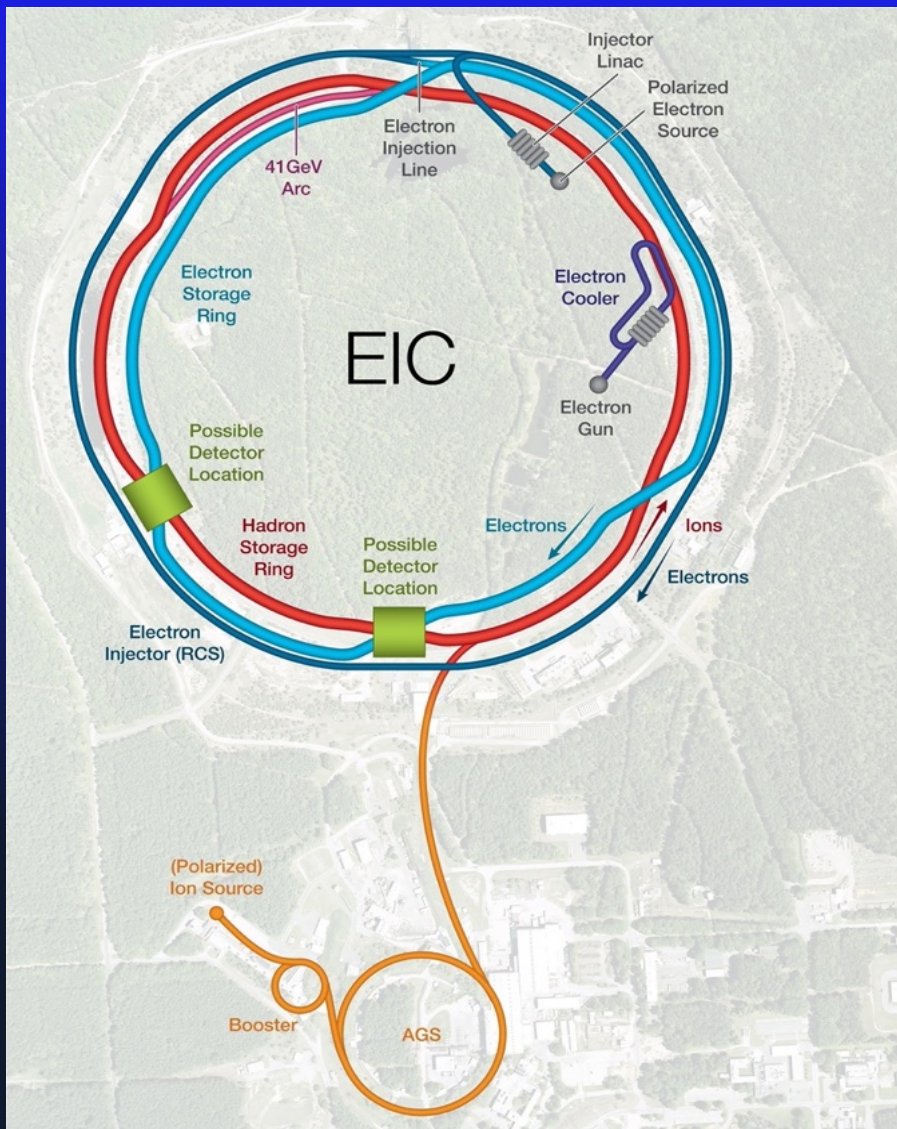


$\text{Au}(197 \times 100 \text{ GeV}) + \text{Au}(197 \times 100 \text{ GeV})$

Brookhaven eRHIC

Add 5-18 GeV electron storage ring in present RHIC tunnel.
Collide e with p to 275 GeV (vs 255 GeV now) and ions to 100 GeV/A in one RHIC ring.

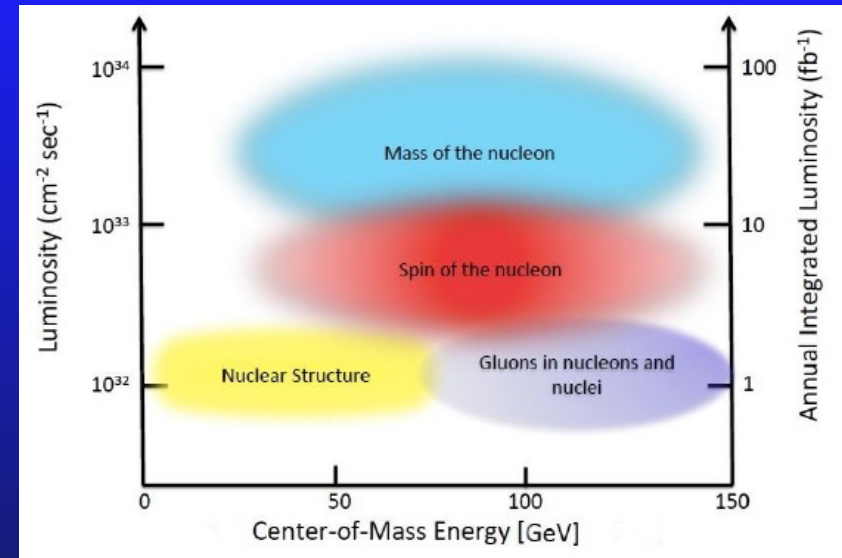
$$\sqrt{s} \simeq 2\sqrt{E_e E_p} = 75-140 \text{ GeV}$$



Accelerator requirements

To map quarks and gluons -- from nuclei (quark-gluon gas) to saturated gluonic matter – designing for variable c.m. energy range from ~28 to ~120 GeV, upgradeable to 140 GeV.

Reach x down to 10^{-4}



Need ion beams from p to heavy stable nuclei: $A = 1 - 208$

3D imaging of gluon and sea quark distributions in nucleons and nuclei requires high luminosity:

L up to 10^{34} /cm² sec. ($\sim 10^2 - 10^3$ X HERA, and LHC at 2×10^{34})

To study correlations of gluon and sea quark distributions with spin, need polarized e^- , p, (and light-ion beams) each above 70%.

Two intersection regions (IP6 and IP8): allow for two detectors

Accelerator challenges

EIC accelerator requirements beyond current technology.
(only large scale accelerator project in the U.S.!)

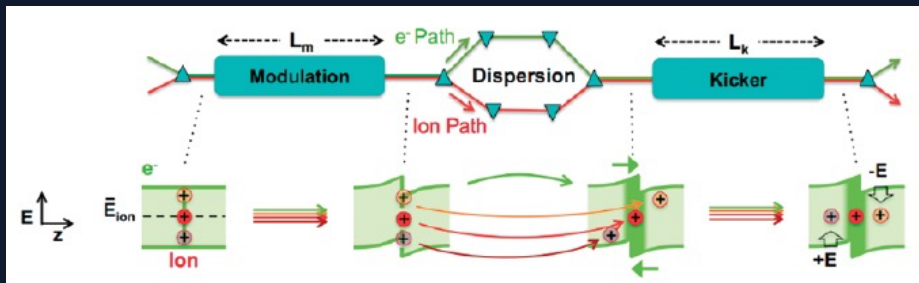
1) High energy, spin-polarized beams colliding with high luminosity.

Polarized beams in a collider achieved only at

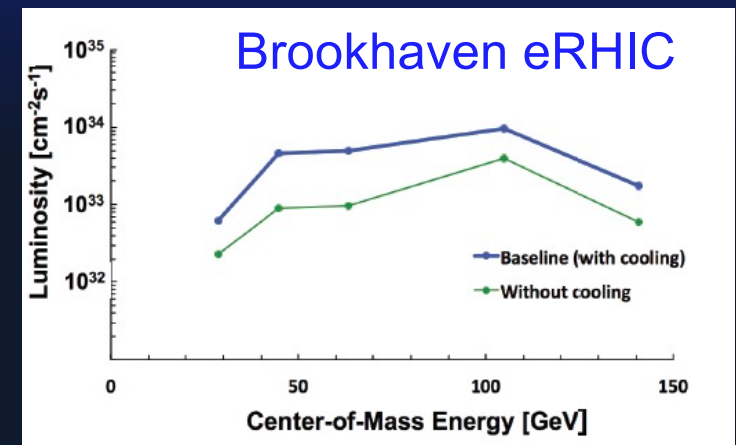
--HERA (polarized e^- or e^+ on unpolarized p) and at the

--Relativistic Heavy Ion Collider (RHIC - pp) with both proton beams polarized

2) Require strong **hadron beam cooling** (focusing) to achieve high luminosity

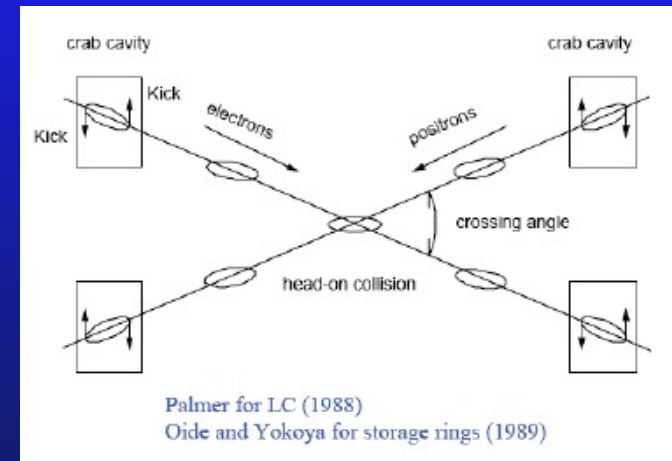
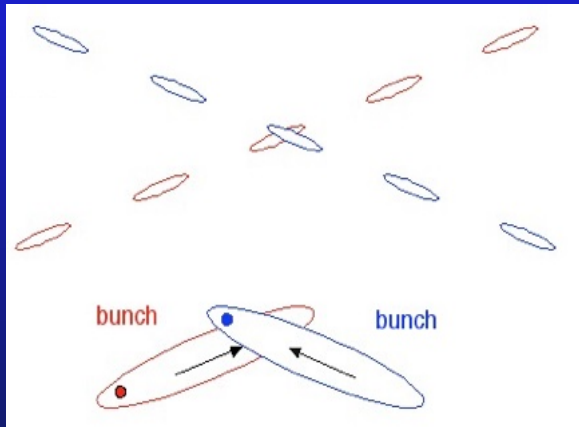


e.g., Coherent Electron Cooling: kick slow hadrons forward, fast ones backward



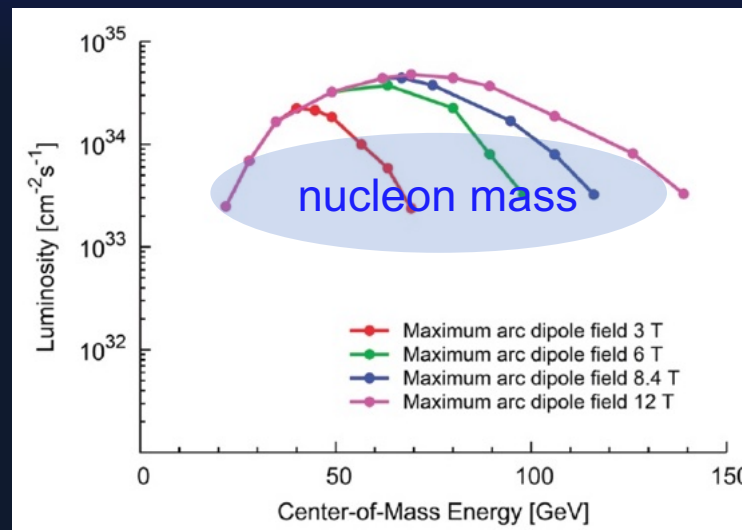
luminosity increase at eRHIC with hadron cooling

3) Intersection region design: Crab crossings (beam walking sideways) to maximize collisions



Instead of bunches crossing at an angle, turn them parallel when they collide parallel. At KEK-B (e^+e^-), but never used in hadron beams.

4) Magnet technology

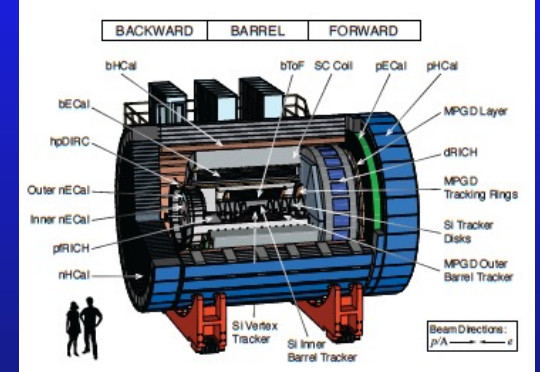


luminosity and
cm e-p energy vs
magnetic fields.
Red curve ~ 10 GeV e^-
+100 GeV p

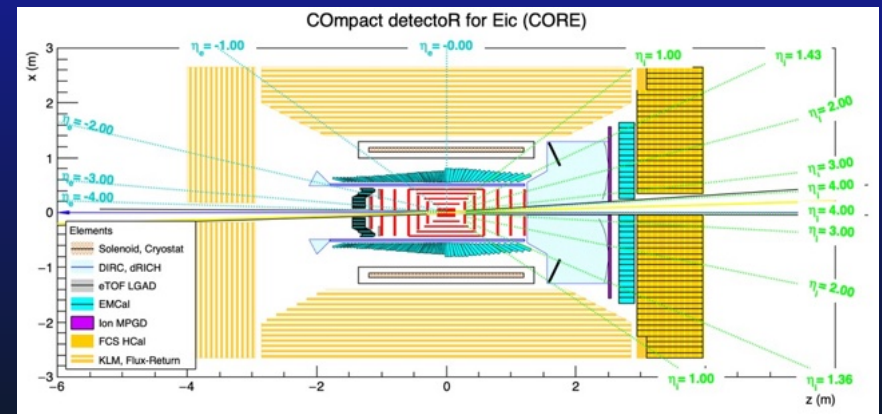
Detectors

Three proposed detectors at IP6:
Possibly a second detector in the future

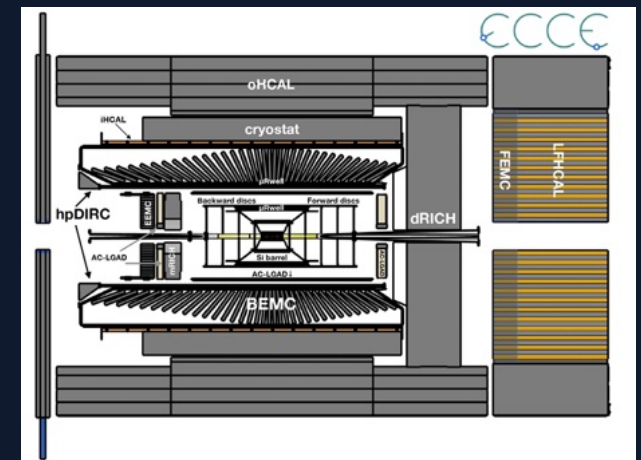
ATHENA (A Totally Hermetic Electron Nucleus Apparatus) *J. Adam et al. J. Instrument. 17, P10019 (2022)*



CORE - a COmpact detectoR for the EIC (arXiv:2209.00496)

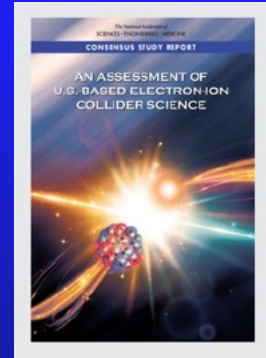


ECCE - EIC Comprehensive Chromodynamics Experiment (arXiv:2209.02580)



EPIC (Electron Proton/Ion Collisions) collaboration finalizing detector design.
Second detector working group formed

Timeline of EIC



Summer 2018: National Academy of Sciences report issued

Sept 2019: EIC enters U.S. budget

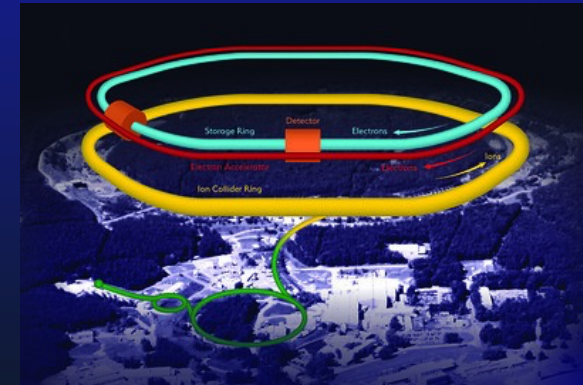
Dec. 19. 2019: Department of Energy first approval (CD-0)

Jan. 9, 2020: Site selection -- Brookhaven

June 2021: Preliminary plan approved (CD-1)

April 2024: CD-3A approved.
"Construction procurement"

April 2025 Approval of final plan (CD-2/3)
(RHIC shutdown June 2025)



Early 2030's: Beams!!
(CD-4)

Department of Energy
(DOE) approval steps

PROJECT ACQUISITION PROCESS AND CRITICAL DECISIONS					
Project Planning Phase		Project Execution Phase			Mission
Preconceptual Planning	Conceptual Design	Preliminary Design	Final Design	Construction	Operations
CD-0	CD-1	CD-2	CD-3	CD-4	
Approve Mission Need	Approve Preliminary Baseline Range	Approve Performance Baseline	Approve Start of Construction	Approve Start of Operations or Project Closeout	



どうもありがとう

